

OPEN 0.26-Hz-linewidth ultrastable lasers at 1557 nm

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Narrow-linewidth ultrastable lasers at 1.5 µm are essential in many applications such as coherent transfer of light through fiber and precision spectroscopy. Those applications all rely on the ultimate performance of the lasers. Here we demonstrate two ultrastable lasers at 1557 nm with a most probable linewidth of 0.26 Hz by independently frequency-stabilizing to the resonance of 10-cm-long ultrastable Fabry-Pérot cavities at room temperature. The fractional frequency instability of each laser system is nearly 8×10^{-16} at 1–30s averaging time, approaching the thermal noise limit of the reference cavities. A remarkable frequency instability of 1×10^{-15} is achieved on the long time scale of 100–4000 s.

Spectrally-narrow laser sources are central to many applications, such as gravitational wave detection, optical atomic clocks, precision spectroscopy and quantum computation¹⁻⁵. To meet those requirements, lasers with a linewidth of a few Hz or sub-hertz have been constructed^{3,6-11}. Among them, many have focused on the particular wavelength of 1.5 µm. Since it is low-loss window of optical fiber, many groups have developed narrow-linewidth ultrastable laser systems at 1.5 µm for coherent transfer of light through fiber 12-15. When compact and low-cost Er:fiber-based laser frequency comb is used to transfer the coherence from 1.5 µm lasers to other optical wavelengths or microwave region 16,17, it enables many novel applications such as direct frequency comb spectroscopy and low noise microwave generation^{4,18}. Furthermore, 1.5 µm narrow-linewidth lasers could also be used to study two-electron interactions via precision spectroscopy of helium¹⁹.

For those attractive applications, narrow-linewidth ultrastable laser systems at 1.5 µm have been constructed based on different techniques. By frequency-stabilizing to the resonance of room-temperature high-finesse ultrastable Fabry-Pérot (F-P) optical cavities with the Pound-Drever-Hall (PDH) technique²⁰, lasers at 1.5 µm with a linewidth of ~1 Hz or sub-hertz have been constructed 14,21. A 40-mHz-linewidth laser has been demonstrated by stabilizing to a silicon single-crystal cavity operated at $124\,\mathrm{K}^{11}$ for low thermal noise 22 . By directly phase-locking to a cavity-stabilized laser in the visible region via an optical frequency comb as a transfer oscillator, a 1 Hz-linewidth laser at 1.5 µm has been constructed²³.

In this paper, the frequencies of two 1557 nm diode lasers are independently stabilized to two 10-cm-long ultrastable F-P cavities with the PDH technique. To achieve a thermal-noise-limited performance, each cavity, operated at room temperature, is deliberately designed for better isolation from environmental vibration and temperature fluctuation. By comparing two similar laser systems, each laser system has a most probable linewidth of $0.26\,\mathrm{Hz}$ and fractional frequency instability of 8×10^{-16} at an averaging time of $1-30\,\mathrm{s}$. Compared to those using cryogenic cavities, the laser systems described in this paper are relatively simple while they still provide excellent frequency stability on both short-term and long-term scale of a few thousand seconds.

Results

Lasers frequency-stabilized to F-P reference cavities. The diagram of the experimental setup of one laser system is shown in Fig. 1. The laser source is a planar external cavity laser operating at 1557 nm with an output power of $10\,\mathrm{mW}$. The output light is firstly frequency-shifted by $+110\,\mathrm{MHz}$ on an acousto-optic modulator (AOM₁), which is used as an isolator and the executor of fast laser frequency control as well. Then the first-order diffracted laser light passes through a combination of a half-waveplate (HWP₁) and a polarization beam splitter (PBS₁), which split the light into two beams with proper light power ratio. One light beam is sent to an F-P reference cavity while the other light beam is used for measurement. In order to isolate the F-P reference cavity from environmental vibration (see Method section), the cavity as well as the optics for the PDH technique are placed on a passive vibration isolation platform and enclosed in an acoustic isolation chamber. A 2-meter-long single-mode (SM) optical fiber is used to deliver the laser light from the laser source to the F-P cavity. Since undesirable random phase noise induced by perturbations via the SM fiber will degrade the light coherence, a fiber noise cancellation (FNC) technique is employed²⁴ (see Method section).

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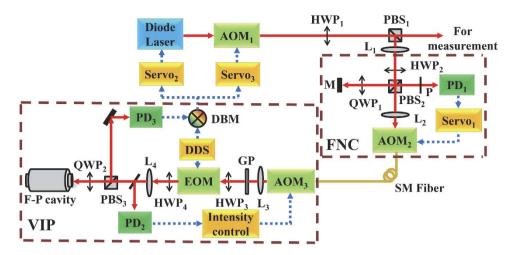


Figure 1. Diagram of the experimental setup. AOM, acousto-optic modulator; HWP, half-wave plate; QWP, quarter-wave plate; L, lens; M, mirror; PBS, polarization beam splitter; P, polarizer; PD, photo-detector; SM fiber, single-mode fiber; EOM, electro-optic modulator; GP, Glan polarizer; DDS, direct digital synthesizer; DBM, double balance mixer; VIP, vibration isolator platform.

The output of the SM optical fiber is firstly frequency-shifted by 110 MHz (+1 order) on AOM₃, whose diffraction efficiency is adjusted by controlling the RF driving power for light intensity stabilization. Then the diffraction light from AOM3 is phase-modulated on an electro-optic modulator (EOM) at the modulation frequency of 50 MHz, where the laser intensity noise approaches shot noise. HWP₃ and a high extinction-ratio Glan polarizer (GP) before the EOM are used to minimize residual amplitude modulation (RAM) of the EOM by adjusting the light polarization to match the main axial of the EO-crystal. Moreover, we controlled the temperature of the EOM slightly above room temperature with a fluctuation of 3-5 mK to reduce RAM which arises from temperature dependent EO-crystal birefringence. After the EOM, a portion of the light is detected on a DC detector (PD₂) to monitor the light power for intensity stabilization. A combination of HWP4 and PBS3 adjust the light power incident to the reference cavity. Lenses L₃ and L₄ are implemented to adjust the size of light beam to match the mode of the F-P cavity. The cavity reflection is picked on PBS₃ and is steered onto a detector (PD₃) by PBS₃. QWP₂ is used to make the light maximally reflected by PBS₃ and prevent feedback of light to the EOM and other optics in the original path. The beat note between the light carrier and modulation sidebands in the cavity reflected light is detected on PD₃. The signal from PD₃ is mixed with part of the driving signal of the EOM on a double balanced mixer (DBM) to obtain a laser frequency discrimination signal. This signal is then sent to Servo2 to control the current injected into the laser module and Servo, to control the driving frequency of AOM₁. In the experimental setup, the laser frequency servo bandwidth is 175 kHz.

Laser linewidth. We have constructed two similar but independent laser systems at 1557 nm and tested the performance of the laser systems by beating against each other on a photodiode. The light from one laser is sent to another system for beating by optical fibers, whose random phase noise is cancelled by active cancellation (see Method section).

To measure the laser linewidth, the beat note between two laser systems is down-converted to $50\,\mathrm{kHz}$ and sent to a spectrum analyzer (Agilent E4402B). As shown in Fig. 2(a), the linewidth of the beat note between two free-running lasers is as large as $11.4\,\mathrm{kHz}$ with a resolution bandwidth (RBW) of $3\,\mathrm{kHz}$ and sweeping measurement time of $70\,\mathrm{ms}$. As long as the laser frequencies are tightly stabilized to the resonance of the ultrastable F-P cavities, the linewidth of the beat note is significantly reduced. For a higher resolution measurement of linewidth, a fast Fourier transform (FFT) spectrum analyzer (SR770) with input frequency range of 0– $100\,\mathrm{kHz}$ is used. As the beat note between the two lasers is $105\,\mathrm{MHz}$, it is down-converted to nearly $1\,\mathrm{kHz}$ and continuously measured by a computer for $10\,\mathrm{hours}$. $1000\,\mathrm{groups}$ of spectra of the beat note were recorded. Each spectrum is measured with a RBW of $122\,\mathrm{mHz}$ and a measurement time of $8\,\mathrm{s}$, which is fitted with a Lorentzian function to obtain the linewidth. A linear frequency drift of $0.03\,\mathrm{Hz/s}$ is removed during measurement by feeding forward to one of synthesizers. Figure 2(b) shows the distribution of the measurement. The most probable linewidth of the beat note between two cavity-stabilized lasers is around $0.36\,\mathrm{Hz}$. The possibility of the linewidth of $0.36\pm0.1\,\mathrm{Hz}$ is 54%. By assuming each laser system contributes equally to the measurement, each laser system has a most probable linewidth of $0.26\,\mathrm{Hz}$.

Laser frequency and phase noise. Besides, the laser frequency noise is measured by stabilizing one of the lasers to Cav₁ and keeping the left portion of the laser light on resonance with Cav₂. Using the PDH signal of Cav₂, the laser frequency noise spectrum is obtained. Figure 2(c) shows the frequency noise spectral density of one laser system (blue solid line). It is close to the calculated thermal noise limit at $0.13 \, \text{Hz}/\sqrt{f}$ (black dash dots) at Fourier frequencies lower than $3 \, \text{Hz}$, where f is the Fourier frequency. Besides, the laser phase noise spectral density $S_{\varphi}(f)$ is shown with red short dash line, which is calculated with the relation of $S_{\psi}(f) = f^2 S_{\varphi}(f)$.

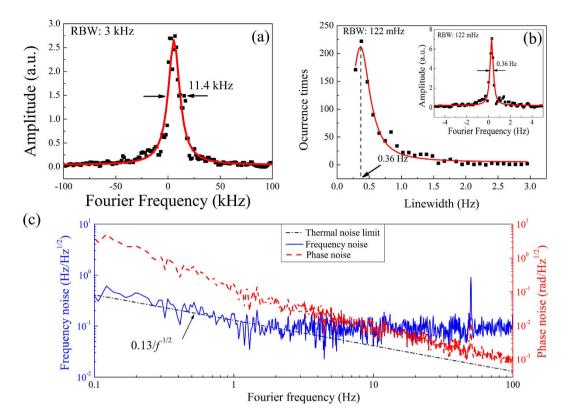


Figure 2. Linewidth measurements, frequency and phase noise spectral density of one laser system. (a) The spectrum of beat note between two free-running lasers (black squares) with 3 kHz RBW and its Lorentzian fitting curve (red solid line). (b) Linewidth distribution of 1000 groups of spectra (black squares) with 122 mHz RBW and Lorentzian fitting curve (red solid line). The inset shows a spectrum of the beat note between two cavity-stabilized lasers (black squares) with 122 mHz RBW and its Lorentzian fitting curve (red solid line). (c) The black dash dots show the calculated thermal-noise-limited frequency noise spectral density of one laser system at $0.13 \text{ Hz}/\sqrt{f}$. The blue solid line shows the frequency noise at Fourier frequencies lower than 3 Hz. The red short dash line shows the calculated laser phase noise based on the relation of $S_{\nu}(f) = f^2 S_{\omega}(f)$.

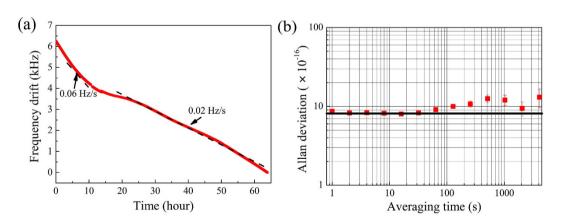


Figure 3. Laser frequency drift and Allan deviation. (a) Relative frequency drift of the beat note between two cavity-stabilized laser systems over 64h (red solid line). The black dash lines show the frequency drift rates. (b) Fractional frequency instability of one cavity-stabilized laser system after a linear drift of $0.03 \, \text{Hz/s}$ is removed. The solid line indicates that the thermal-noise-limited frequency instability of 8×10^{-16} .

Frequency instability. The relative frequency drift between the two cavity-stabilized systems is measured by counting the beating frequency of two lasers over 64 hours (The counter is Agilent 53132A). As shown in Fig. 3(a), the total frequency drift is 6.25 kHz. The frequency drift rate is within 0.08 Hz/s. In 90% of time, the laser drifts less than 0.04 Hz/s. The fractional frequency instability of one cavity-stabilized laser is shown in Fig. 3(b) after

Figure 4. Shape of cavity and calculation results. (a) Three-dimensional view of the reference cavity and its support. The cavity is placed on a U-shape block and supported by four hemisphere-rubbers which are 30 mm away from the ends of the cylinder spacer along the optical axis. **(b)** The vibration sensitivity of probe points on mirrors of the reference cavity along *z* axis to accelerations along *z* axis when *d* varies according to finite element analysis. Probe points at the position of 0 mm are at the center of the mirrors.

removing a linear drift of 0.03 Hz/s during the calculation. At an averaging time of 1–30 s, the fractional frequency instability of one laser (Allan deviation) reaches the thermal-noise-limited performance of the reference cavity of 8×10^{-16} . Benefiting from the thermal shields inserted between the FP cavity and the vacuum chamber, the laser frequency instability at an averaging time of 100-4000 s can be as low as 1×10^{-15} .

Discussion

In conclusion, two lasers operating at 1557 nm are independently stabilized to two F-P cavities. With good isolation of the F-P reference cavities from environmental perturbations, the reference cavities approach a thermal-noise-limited performance. The most probable laser linewidth is 0.26 Hz and the laser frequency instability is around 8×10^{-16} at the 1–30 s averaging time. For a time scale of 100–4000 s, laser frequency instability of 1×10^{-15} is achieved.

The laser systems reported in this paper is limited by the thermal noise of the reference cavities. To further improve the performance of the lasers, it is inevitable to reduce the thermal noise of the reference cavities. Either using cryogenic cavities with the coefficient of thermal expansion close to zero at low temperatures or using room-temperature cavities with longer length 7,11,25 , the thermal-noise-limited laser frequency instability will be reduced to 1×10^{-16} . However, to keep the laser system as simple as possible, it might be a solution by using low mechanical loss mirrors and coatings 8,25,26 . For a 10-cm-long cavity with fused silica (FS) mirrors coated with crystalline coating of AlGaAs, the thermal-noise-limited laser frequency instability will approach 3×10^{-16} .

We are transferring the coherence of the 0.26 Hz-linewidth laser light to remote locations through fiber links for potential applications such as precision spectroscopy and low noise microwave generation at remote ends.

Methods

Fiber noise cancellation. In the FNC, the fiber phase noise is detected by comparing the fiber round-trip light against the local reference light on a Michelson interferometer. As shown in Fig. 1, a combination of HWP₂ and PBS₂ splits light into two interferometric arms with suitable intensity ratio. The light in the short reference arm goes through a quarter-wave plate (QWP_1) and is reflected by a high reflectance mirror. Then it passes through QWP_1 again and goes through PBS₂. The light in the long arm is frequency-shifted by +77.5 MHz on AOM_2 and is coupled into the SM optical fiber. At the remote end of the SM fiber, a small portion of the light is reflected by the flat surface of the SM fiber at the remote end and passes through AOM_2 again. Both light beams from the two arms are combined on PBS₂ and beat against each other on a photo detector (PD_1) . The beating signal of 155 MHz is mixed with a 155 MHz local oscillator to extract the fiber phase noise. This signal is used as an error signal to compensate the fiber phase noise by tuning the driving frequency of AOM_2 . In order to protect the optics from air flow, both the laser and optics for the FNC system are sealed in an acrylic box.

F-P cavities. The dominant limitation of the laser frequency instability is usually the length instability of F-P cavity.

To reduce the thermal expansion of the F-P cavities, each F-P cavity is made of ultralow expansion (ULE) glass, whose coefficient of thermal expansion (CTE) is within ± 30 ppb/K. Two mirrors made of ULE glass with curvature of $R_1 = \infty$ and $R_2 = 0.5$ m are optically-contacted to a 10-cm-long cylinder spacer. To reduce the thermal sensitivity, each cavity is located in two layers of thermal shields as low-pass thermal filters. The inner layer is made of gold-plated copper while the outer layer is made of high-polished aluminum for less weight. The reference cavity as well as two layers of thermal shields is enclosed in an aluminum vacuum chamber evacuated to 5×10^{-7} torr. This kind of configuration is calculated to have a thermal time constant of 28.8 hours²⁷, which helps to reduce the thermal sensitivity of the reference cavity to environmental thermal fluctuation with a period of 1 hour by three orders of magnitude when compared with the system without any thermal shield. Then the vacuum chamber is enclosed in thermal isolations and temperature-stabilized above room temperature with a

fluctuation of < 2 mK. The vacuum chamber and the optics for the PDH technique are on the vibration isolator platform, which resides in a box made of aluminum alloy plates covered with acoustic absorption foam. The box protects the cavity from air flow, temperature fluctuation and acoustic noise. Two F-P reference cavities are independently placed on two passive vibration isolator platforms and enclosed in two acoustic isolation chambers.

In order to reduce the vibration sensitivity of the F-P cavities, the cavity supporting configuration is designed according to finite element analysis with ANSYS. Figure 4(a) shows the three-dimensional view of the cavity. The F-P cavity is horizontally mounted on four hemisphere rubbers, which are sitting on a U-shape fused silica block. Figure 4(b) shows the simulated sensitivity of the cavity length to an acceleration of $g = 9.8 \text{ m/s}^2$ along z axis. When the hemisphere rubbers are 30 mm away from the ends of the cylinder spacer along the optical axis (x axis), the cavity has the smallest vibration sensitivity of $5.5 \times 10^{-13}/g$ at the center of the mirror.

The finesse of the cavities is measured to be \sim 270,000 for Cav₁ and \sim 250,000 for Cav₂ from frequency-sweep optical-heterodyne cavity-ring-down spectroscopy²⁸, corresponding to a cavity linewidth of \sim 5.6 kHz and \sim 6 kHz, respectively. The cavity reflection contrasts are 67% for Cav₁ and 60% for Cav₂, respectively.

References

- 1. Waldman, S. J. Status of LIGO at the start of the fifth science run. Class. Quantum Grav. 23, S653-S660 (2006).
- 2. Ludlow, A. D., Boyd, M. M., Ye, J., Peik, E. & Schmidt, P. O. Optical atomic clocks. Rev. Mod. Phys. 87, 637–701 (2015).
- 3. Jiang, Y. Y. et al. Making optical atomic clocks more stable with 10⁻¹⁶-level laser stabilization. Nat. Photonics 5, 158–161 (2011).
- 4. Rieker, G. B. et al. Frequency-comb-based remote sensing of greenhouse gases over kilometer air paths. Optica 1, 290 (2014).
- 5. Leibfried, D., Blatt, R., Monroe, C. & Wineland, D. Quantum dynamics of single trapped ions. Rev. Mod. Phys. 75, 281–324 (2003).
- 6. Webster, S. A., Oxborrow, M. & Gill, P. Subhertz-linewidth Nd:YAG laser. Opt. Lett. 29, 1497-1499 (2004).
- 7. Häfner, S. et al. 8×10^{-17} fractional laser frequency instability with a long room-temperature cavity. Opt. Lett. **40**, 2112–2115 (2015).
- 8. Ludlow, A. D. et al. Compact, thermal-noise-limited optical cavity for diode laser stabilization at 1×10^{-15} . Opt. Lett. 32, 641–643 (2007).
- 9. Alnis, J. et al. Subhertz linewidth diode lasers by stabilization to vibrationally and thermally compensated ultralow-expansion glass Fabry-Perot cavities. Phys. Rev. A 77, 053809 (2008).
- Chen, H., Jiang, Y., Fang, S., Bi, Z. & Ma, L. Frequency stabilization of Nd:YAG lasers with a most probable linewidth of 0.6 Hz. J. Opt. Soc. Am. B 30, 1546–1550 (2013).
- 11. Kessler, T. et al. A sub-40 mHz-linewidth laser based on a silicon single-crystal optical cavity. Nat. Photonics 6, 687-692 (2012).
- 12. Predehl, K. et al. A 920-kilometer optical fiber link for frequency metrology at the 19th decimal place. Science 336, 441–444 (2012).
- 13. Argence, B. et al. Quantum cascade laser frequency stabilization at the sub-Hz level. Nat. Photonics 9, 456-460 (2015).
- 14. Newbury, N. R., Williams, P. A. & Swann, W. C. Coherent transfer of an optical carrier over 251 km. *Opt. Lett.* **32**, 3056–3058 (2007).
- 15. Musha, M., Hong, F. L., Nakagawa, K. & Ueda, K. I. Coherent optical frequency transfer over 50-km physical distance using a 120-km-long installed telecom fiber network. *Opt. Express* 16, 16459–16466 (2008).
- 16. Hagemann, C. et al. Providing 10⁻¹⁶ short-term stability of a 1.5-μm laser to optical clocks. IEEE Trans. Instrum. Meas. 62, 1556 (2013).
- 17. Fortier, T. M. et al. Generation of ultrastable microwaves via optical frequency division. Nat. Photonics 5, 425-429 (2011).
- 18. Quinlan, F. et al. Ultralow phase noise microwave generation with an Er:fiber-based optical frequency divider. Opt. Lett. 36, 3260–3262 (2011).
- 19. Van Rooij, R. *et al*. Frequency metrology in quantum degenerate helium: direct measurement of the $2^{3}S_{1} \rightarrow 2^{1}S_{0}$ transition. *Science* 333, 196–198 (2013).
- 20. Drever, R. W. P. *et al.* Laser phase and frequency stabilization using an optical resonator. *Appl. Phys. B* **31**, 97–105 (1983).
- 21. Jiang, H. et al. Long-distance frequency transfer over an urban fiber link using optical phase stabilization. J. Opt. Soc. Am. B 25, 2029–2035 (2008).
- 22. Numata, K., Kemery, A. & Camp, J. Thermal-noise limit in the frequency stabilization of lasers with rigid caviteis. *Phys. Rev. Lett.* **93**, 250602 (2004).
- 23. Grosche, G. et al. Optical frequency transfer via $146 \,\mathrm{km}$ fiber link with 10^{-19} relative accuracy. Opt. Lett. 34, 2270–2272 (2009).
- 24. Ma, L. S., Jungner, P., Ye, J. & Hall, J. L. Delivering the same optical frequency at two places: accurate cancellation of phase noise introduced by an optical fiber or other time-varying path. *Opt. Lett.* 19, 1777–1779 (1994).
- 25. Notcutt, M., Ma, L. S., Ludlow, A. D., Foreman, S. M., Ye, J. & Hall, J. L. Contribution of thermal noise to frequency stability of rigid optical cavity via Hertz-linewidth lasers. *Phys. Rev. A* 73, 031804R (2006).
- 26. Cole, G. D., Zhang, W., Martin, M. J., Ye, J. & Aspelmeyer, M. Tenfold reduction of Brownian noise in high-reflectivity optical coatings. *Nat. Photonics* 7, 644–650 (2013).
- 27. Dai, X., Jiang, Y., Hang, C., Bi, Z. & Ma, L. Thermal analysis of optical reference cavities for low sensitivity to environmental temperature fluctuations. *Opt. Express* 23, 5134–5146 (2015).
- 28. Ye, J. "Ultrasensitive high resolution laser spectroscopy and its application to optical frequency standards," Ph.D. thesis (University of Colorado at Boulder, 1997).

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Author Contributions

Y.J. and L.M. designed the experiment; L.W. and C.M. constructed the systems and measured the experimental data; W.Q. performed the simulation on the vibration sensitivity of F-P reference cavities; H.Y. and Z.B. prepared electrical circuits; all authors contributed to the manuscript.

Additional Information

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