



OPEN

Machine learning and serving of discrete field theories

Hong Qin

A method for machine learning and serving of discrete field theories in physics is developed. The learning algorithm trains a discrete field theory from a set of observational data on a spacetime lattice, and the serving algorithm uses the learned discrete field theory to predict new observations of the field for new boundary and initial conditions. The approach of learning discrete field theories overcomes the difficulties associated with learning continuous theories by artificial intelligence. The serving algorithm of discrete field theories belongs to the family of structure-preserving geometric algorithms, which have been proven to be superior to the conventional algorithms based on discretization of differential equations. The effectiveness of the method and algorithms developed is demonstrated using the examples of nonlinear oscillations and the Kepler problem. In particular, the learning algorithm learns a discrete field theory from a set of data of planetary orbits similar to what Kepler inherited from Tycho Brahe in 1601, and the serving algorithm correctly predicts other planetary orbits, including parabolic and hyperbolic escaping orbits, of the solar system without learning or knowing Newton's laws of motion and universal gravitation. The proposed algorithms are expected to be applicable when the effects of special relativity and general relativity are important.

Data-driven methodology has attracted much attention recently in the physics community. This is not surprising since one of the fundamental objectives of physics is to deduce or discover the laws of physics from observational data. The rapid development of artificial intelligence technology begs the question of whether such deductions or discoveries can be carried out algorithmically by computers.

In this paper, I propose a method for machine learning of discrete field theories in physics from observational data. The method also includes an effective algorithm to serve the discrete field theories learned, in terms of predicting new observations.

Machine learning is not exactly a new concept in physics. In particular, the connection between artificial neural networks and dynamical systems has been noticed for decades^{1–8}. What is the new contribution brought by the present study? Most current applications of machine learning techniques in physics roughly fall into the following categories. (i) Using neural networks to model complex physical processes, such as plasma disruptions in magnetic fusion devices^{9–12}, effective Reynolds stress due to turbulence¹³, coarse-grained nonlinear effects¹⁴, and proper moment closure schemes for fluid systems¹⁵. (ii) Solving differential equations in mathematical physics by approximating solutions with neural networks^{16–21}. In particular, significant progress has been made in solving Schrödinger's equation for many-body systems^{22,23}. (iii) Discovering unknown functions or undetermined parameters in governing differential equations^{24–33}. As a specific example, methods of learning the Hamiltonian function of a canonical symplectic Hamiltonian system were proposed very recently^{34–41}. (iv) Using neural networks to generate sampling data in statistical ensembles for calculating equilibrium properties of physical systems^{42–45}.

The problem addressed in this paper belongs to a new category. The method proposed learns a field theory from a given set of training data consisting of observed values of a physical field at discrete spacetime locations. The laws of physics are fundamentally expressed in the form of field theories instead of differential equations. It is thus more important to learn the underpinning field theories when possible. Since field theories are in general simpler than the corresponding differential equations, learning field theories is easier, which is true for both human intelligence and artificial intelligence. Except for the fundamental assumption that the observational data are governed by field theories, the learning and serving algorithms proposed do not assume any knowledge of the laws of physics, such as Newton's law of motion and Schrödinger's equation. This is different from most existing methodologies of machine learning in physics.

Without loss of generality, let's briefly review the basics of field theories using the example of first-order field theory in the space of \mathbb{R}^n for a scalar field ψ . A field theory is specified by a Lagrangian density $L(\psi, \partial\psi/\partial x^\alpha)$, where x^α ($\alpha = 1, \dots, n$) are the coordinates for \mathbb{R}^n . The theory requires that with the value of ψ fixed at the boundary, $\psi(x)$ varies with respect to x in such a way that the action of the system

Plasma Physics Laboratory, Princeton University, Princeton, NJ 08543, USA. email: hongqin@princeton.edu

$$\mathcal{A} = \int L(\psi, \partial\psi/\partial x^\alpha) d^n x \quad (1)$$

is minimized. Such a requirement of minimization is equivalent to the condition that the following Euler-Lagrange (EL) equation is satisfied everywhere in \mathbb{R}^n ,

$$EL(\psi) \equiv \sum_{\alpha=1}^n \frac{\partial}{\partial x^\alpha} \left(\frac{\partial L}{\partial(\partial\psi/\partial x^\alpha)} \right) - \frac{\partial L}{\partial\psi} = 0. \quad (2)$$

The problem of machine learning of field theories can be stated as follows:

Problem Statement 1. For a given set of observed values of ψ on a set of discrete points in \mathbb{R}^n , find the Lagrangian density $L(\psi, \partial\psi/\partial x^\alpha)$ as a function of ψ and $\partial\psi/\partial x^\alpha$, and design an algorithm to predict new observations of ψ from L .

Now it is clear that learning the Lagrangian density $L(\psi, \partial\psi/\partial x^\alpha)$ is easier than learning the EL equation (2), which depends on ψ in a more complicated manner than L does. For example, the EL equation depends on second-order derivatives $\partial^2\psi/\partial x^\alpha\partial x^\beta$ and L does not. However, learning L from a given set of observed values of ψ is not an easy task either for two reasons. Suppose that L is modeled by a neural network. We need to train L using the EL equation, which requires the knowledge of $\partial^2\psi/\partial x^\alpha\partial x^\beta$. For this purpose, we can set up another neural network for $\psi(x)$, which needs to be trained simultaneously with L . This is obviously a complicated situation. Alternatively, one may wish to calculate $\partial^2\psi/\partial x^\alpha\partial x^\beta$ from the training data. But it may not be possible to calculate them with desired accuracy, depending on the nature of the training data. Secondly, even if the optimized neural network for L is known, serving the learned field theory by solving the EL equation with a new set of boundary conditions presents a new challenge. The first-order derivatives $\partial\psi/\partial x^\alpha$ and second-order derivatives $\partial^2\psi/\partial x^\alpha\partial x^\beta$ are hidden inside the neural network for L , which is nonlinear and possibly deep. Solving differential equations defined by neural networks ventures into uncharted territory.

As will be shown in “Machine learning and serving of discrete field theories” section, reformulating the problem in terms of discrete field theory overcomes both difficulties. Problem Statement 1 will be replaced by Problem Statement 2 in “Machine learning and serving of discrete field theories” section. To learn a discrete field theory, it suffices to learn a discrete Lagrangian density L_d , a function with $n + 1$ inputs, which are the values of ψ at $n + 1$ adjacent spacetime locations. The training of L_d is straightforward. Learning serves the purpose of serving, and the most effective way to serve a field theory with long term accuracy and fidelity is by offering the discrete version of the theory, as has been proven by the recent advances in structure-preserving geometric algorithms^{46–75}. Therefore, learning a discrete field theory directly from the training data and then serving it constitute an attractive approach for discovering physical models by artificial intelligence.

It has long been theorized since Euclid’s study on mirrors and optics that as the most fundamental law of physics, all nature does is to minimize certain actions^{76,77}. But how does nature do that? The machine learning and serving algorithms of discrete field theories proposed might provide a clue, when incorporating the basic concept of the simulation hypothesis by Bostrom⁷⁸. The simulation hypothesis states that the physical universe is a computer simulation, and it is being carefully examined by physicists as a possible reality^{79–81}. If the hypothesis is true, then the spacetime is necessarily discrete. So are the field theories in physics. It is then reasonable, at least from a theoretical point of view, to suggest that some machine learning and serving algorithms of discrete field theories are what the discrete universe, i.e., the computer simulation, runs to minimize the actions.

In “Machine learning and serving of discrete field theories” section, the learning and serving algorithms of discrete field theories are developed. Two examples of learning and predicting nonlinear oscillations in 1D are given in “Examples of learning and predicting nonlinear oscillations” section to demonstrate the method and algorithms. In “Kepler problem” section, I apply the methodology to the Kepler problem. The learning algorithm learns a discrete field theory from a set of observational data for orbits of the Mercury, Venus, Earth, Mars, Ceres, and Jupiter, and the serving algorithm correctly predicts other planetary orbits, including the parabolic and hyperbolic escaping orbits, of the solar system. It is worthwhile to emphasize that the serving and learning algorithms do not know, learn, or use Newton’s laws of motion and universal gravitation. The discrete field theory directly connect the observational data and new predictions. Newton’s laws are not needed.

Machine learning and serving of discrete field theories

In this section, I describe first the formalism of discrete field theory on a spacetime lattice, and then the algorithm for learning discrete field theories from training data and the serving algorithm to predict new observations using the learned discrete field theories. The connection between the serving algorithm and structure-preserving geometric integration methods is highlighted.

To simplify the presentation and without losing generality, the theory and algorithms are given for the example of a first-order scalar field theory in \mathbb{R}^2 . One of the dimension will be referred to as time with coordinate t , and the other dimension space with coordinate x . Generalizations to high-order theories and to tensor fields or spinor fields are straightforward.

For a discrete field theory in \mathbb{R}^2 , the field $\psi_{i,j}$ is defined on a spacetime lattice labeled by two integer indices (i, j) . For simplicity, let’s adopt a rectangular lattice shown in Fig. 1. The first index i identifies temporal grid points, and the second index j spacial grid points. The discrete action \mathcal{A}_d of the system is the summation of discrete Lagrangian densities over all grid cells,

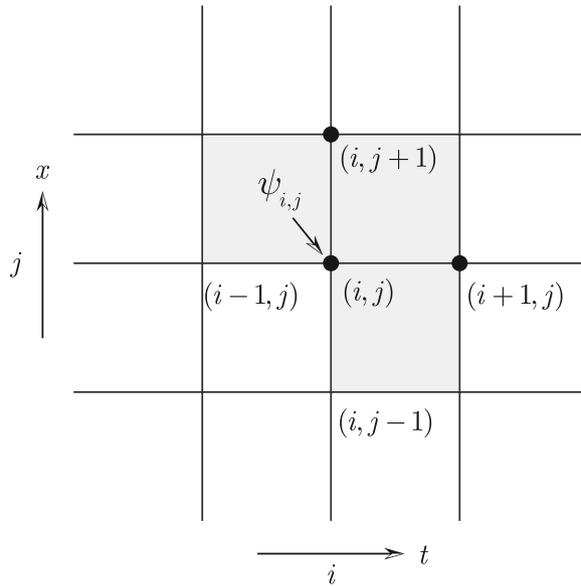


Figure 1. Spacetime lattice and discrete field ψ . The discrete Lagrangian density $L_d(\psi_{i,j}, \psi_{i+1,j}, \psi_{i,j+1})$ of the grid cell whose lower left vertex is at the grid point (i, j) is chosen to be a function of the values of the discrete field at the three vertices marked by solid circles. The action \mathcal{A}_d of the system depends on $\psi_{i,j}$ through the discrete Lagrangian densities of the three neighboring grid cells indicated by gray shading.

$$\mathcal{A}_d = \Delta t \Delta x \sum_{i,j} L_d(\psi_{i,j}, \psi_{i+1,j}, \psi_{i,j+1}), \tag{3}$$

where Δt and Δx are the grid sizes in time and space respectively, and $L_d(\psi_{i,j}, \psi_{i+1,j}, \psi_{i,j+1})$ is the discrete Lagrangian density of the grid cell whose lower left vertex is at the grid point (i, j) . I have chosen L_d to be a function of $\psi_{i,j}$, $\psi_{i+1,j}$, and $\psi_{i,j+1}$ only, which is suitable for first-order field theories. For instance, in the continuous theory for wave dynamics, the Lagrangian density is

$$L\left(\psi, \frac{\partial \psi}{\partial t}, \frac{\partial \psi}{\partial x}\right) = \left(\frac{\partial \psi}{\partial t}\right)^2 - \left(\frac{\partial \psi}{\partial x}\right)^2. \tag{4}$$

Its counterpart in the discrete theory can be written as

$$L_d(\psi_{i,j}, \psi_{i+1,j}, \psi_{i,j+1}) = \left(\frac{\psi_{i+1,j} - \psi_{i,j}}{\Delta t}\right)^2 - \left(\frac{\psi_{i,j+1} - \psi_{i,j}}{\Delta x}\right)^2. \tag{5}$$

The discrete Lagrangian density L_d defined in Eq. (5) can be viewed as an approximation of the continuous Lagrangian density L in Eq. (4). But I prefer to take L_d as an independent object that defines a discrete field theory.

For the discrete field theory, the condition of minimizing the discrete action \mathcal{A}_d with respect to each $\psi_{i,j}$ demands

$$EL_{i,j}(\psi) \equiv \frac{\partial \mathcal{A}_d}{\partial \psi_{i,j}} = \frac{\partial}{\partial \psi_{i,j}} [L_d(\psi_{i-1,j}, \psi_{i,j}, \psi_{i-1,j+1}) + L_d(\psi_{i,j}, \psi_{i+1,j}, \psi_{i,j+1}) + L_d(\psi_{i,j-1}, \psi_{i+1,j-1}, \psi_{i,j})] = 0. \tag{6}$$

Equation (6) is called Discrete Euler-Lagrange (DEL) equation for the obvious reason that its continuous counterpart is the EL equation (2) with $x^1 = t$ and $x^2 = x$. Following the notation of the continuous theory, I also denote the left-hand-side of the last equal sign in Eq. (6) by an operator $EL_{i,j}(\psi)$, which maps the discrete field $\psi_{i,j}$ into another discrete field. The DEL equation is employed to solve for the discrete field ψ on the spacetime lattice when a discrete Lagrangian density L_d is prescribed. This has been the only usage of the DEL equation in the literature that I am aware of so far^{48,49,51–53,55,57,58,61,64,68,70–75}. I will come back to this shortly.

For the problem posed in the present study, the discrete Lagrangian density L_d is unknown. It needs to be learned from the training data. Specifically, in terms of the discrete field theory, the learning problem discussed in “Introduction and statement of the problem” section becomes:

Problem Statement 2. For a given set of observed data $\bar{\psi}_{i,j}$ on a spacetime lattice, find the discrete Lagrangian density $L_d(\psi_{i,j}, \psi_{i+1,j}, \psi_{i,j+1})$ as a function of $\psi_{i,j}$, $\psi_{i+1,j}$, and $\psi_{i,j+1}$, and design an algorithm to predict new observations of $\psi_{i,j}$ from L_d .

Unlike the difficult situation described in “Introduction and statement of the problem” section for learning a continuous field theory, learning a discrete field theory is straightforward. The algorithm is obvious once the problem is declared as in Problem Statement 2. We set up a function approximation for L_d with three inputs and one output using a neural network or any other approximation scheme adequate for the problem under investigation. The approximation is optimized by adjusting its free parameters to minimize the loss function

$$F(\bar{\psi}) = \frac{1}{IJ} \sum_{i=1}^{I-1} \sum_{j=1}^{J-1} EL_{i,j}(\bar{\psi})^2 \quad (7)$$

on the training data $\bar{\psi}$, where I and J are the total number of grid points in time and space respectively. In Problem Statement 2 and the definition of loss function (7), it is implicitly assumed that the training data are available over the entire spacetime lattice. Notice that according to Eqs. (6) and (7), first-order derivatives of L_d with respect to all three arguments are required to evaluate $F(\bar{\psi})$. Automatic differential algorithms²⁹, which have been widely adopted in artificial neural networks, can be applied. To train the neural network or other approximation for L_d , established methods, including Newton’s root searching algorithm and the Adam optimizer³², are available.

Once the discrete Lagrangian density L_d is trained, the learned discrete field theory is ready to be served to predict new observations. After boundary conditions are specified, the DEL equation (6) is solved for the discrete field $\psi_{i,j}$. A first-order field theory requires two boundary conditions in each dimension. As an illustrative example, let’s assume that $\psi_{0,j}$ and $\psi_{1,j}$ are specified for all j s, corresponding to two initial conditions at $t = 0$, and $\psi_{i,0}$ and $\psi_{i,1}$ are specified for all i s, corresponding to two boundary conditions at $x = 0$. Under these boundary and initial conditions, the DEL equation (6) can be solved for field $\psi_{i,j}$ for all i s and j s as follows.

- Step 1: Start from the DEL equation at $(i, j) = (1, 2)$, i.e., $EL_{1,2}(\psi) = 0$, which is an algebraic equation containing only one unknown $\psi_{2,2}$. Solve $EL_{1,2}(\psi) = 0$ for $\psi_{2,2}$ using a root searching algorithm, e.g., Newton’s algorithm.
- Step 2: Move to grid point $(i, j) = (1, 3)$. Solve the DEL equation $EL_{1,3}(\psi) = 0$ for the only unknown $\psi_{2,3}$.
- Step 3: Repeat Step 2) with increasing value of j to generate solution $\psi_{2,j}$ for all j s.
- Step 4: Increase index i to 2. Apply the same procedure in Step 3) for generating $\psi_{2,j}$ to generate $\psi_{3,j}$ for all j s.
- Step 5: Repeat Step 4) for $i = 3, 4, \dots, I$ to solve for all $\psi_{i,j}$.

In a nutshell, the DEL equation at the grid cell labeled by (i, j) (see Fig. 1) is solved as an algebraic equation for $\psi_{i+1,j}$. This serving algorithm propagates the solution from the initial and boundary conditions to the entire spacetime lattice. It is exactly how the physical field propagates physically. According to the simulation hypothesis, the algorithmic propagation and the physical propagation are actually the same thing. When different types of boundary and initial conditions are imposed, the algorithm needs to be modified accordingly. But the basic strategy remains the same. Specific cases will be addressed in future study.

The above algorithms in R^2 can be straightforwardly generalized to R^n , where the discrete Lagrangian density L_d will be a function of $n + 1$ variables, i.e., $\psi_{i_1, i_2, \dots, i_n}, \psi_{i_1+1, i_2, \dots, i_n}, \psi_{i_1, i_2+1, \dots, i_n}, \dots, \psi_{i_1, i_2, \dots, i_n+1}$. And in a similar way as in R^2 , the serving algorithm solves for $\psi_{i_1, i_2, \dots, i_n}$ by propagating its values at the boundaries to the entire lattice. It can also be easily generalized to vector fields or spinor fields, as exemplified in “Kepler problem” section.

It turns out this algorithm to serve the learned discrete field theory is a variational integrator. The principle of variational integrators is to discretize the action and Lagrangian density instead of the associated EL equations. Methods and techniques of variational integrators have been systematically developed in the past decade^{46–75}. The advantages of variational integrators over standard integration schemes based on discretization of differential equations have been demonstrated. For example, variational integrators in general are symplectic or multi-symplectic^{48,49,51–53,55,57,58,61,64,68,70–75}, and as such are able to bound globally errors on energy and other invariants of the system for all simulation time-steps. More sophisticated discrete field theories have been designed to preserve other geometric structures of physical systems, such as the gauge symmetry^{52,75} and Poincaré symmetry^{72,80,81,83}. What proposed in this paper is to learn the discrete field theory directly from observational data and then serve the learned discrete field theory to predict new observations.

Examples of learning and predicting nonlinear oscillations

In this section, I use two examples of learning and predicting nonlinear oscillations in 1D to demonstrate the effectiveness of the learning and serving algorithms. In 1D, the discrete action reduces to the summation of the discrete Lagrangian density L_d over the time grids,

$$\mathcal{A}_d = \Delta t \sum_{i=0}^I L_d(\psi_i, \psi_{i+1}). \quad (8)$$

Here, $L_d(\psi_i, \psi_{i+1})$ is a function of the field at two adjacent time grid points. The DEL equation is simplified to

$$EL_i(\psi) \equiv \frac{\partial L_d(\psi_{i-1}, \psi_i)}{\partial \psi_i} + \frac{\partial L_d(\psi_i, \psi_{i+1})}{\partial \psi_i} = 0. \quad (9)$$

The training data $\bar{\psi}_i (i = 0, \dots, I)$ form a time sequence, and the loss function on a data set ψ is

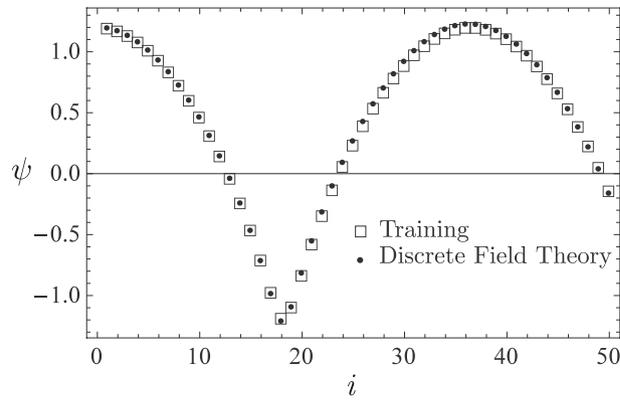


Figure 2. The predicted sequence ψ_i (solid circles) from the learned discrete field theory and the training sequence $\bar{\psi}_i$ (empty squares) are barely distinguishable in the figure. The discrete Lagrangian is trained until the loss function $F(\bar{\psi})$ is less than 10^{-7} .

$$F(\bar{\psi}) = \frac{1}{I} \sum_{i=1}^{I-1} EL_i(\bar{\psi})^2,$$

After learning L_d , the serving algorithm will predict a new time sequence for every two initial conditions ψ_0 and ψ_1 . Note that Eq. (9) is an algebraic equation for ψ_{i+1} when ψ_{i-1} and ψ_i are known. It is an implicit two-step algorithm from the viewpoint of numerical methods for ordinary differential equations. It can be proven^{49,51} that the algorithm exactly preserves a symplectic structure defined by

$$\Omega(\psi_i, \psi_{i+1}) = d\theta, \tag{10}$$

$$\theta = \frac{\partial L_d(\psi_i, \psi_{i+1})}{\partial \psi_{i+1}} d\psi_{i+1}. \tag{11}$$

The algorithm is thus a symplectic integrator, which is able to bound globally the numerical error on energy for all simulation time-steps. Compared with standard integrators which do not possess structure-preserving properties, such as the Runge-Kutta method, variational integrators deliver much improved long-term accuracy and fidelity.

For each of the two examples, the training data taken by the learning algorithm are a discrete time sequence generated by solving the EL equation of an exact continuous Lagrangian. In 1D, the EL equation is an Ordinary Differential Equation (ODE) in time. Only the training sequence is visible to the learning and serving algorithms, and the EL equation and the continuous Lagrangian are not. After learning the discrete Lagrangian from the training data, the algorithm serves it by predicting new dynamic sequences ψ_i for different initial conditions. The predictions are compared with accurate numerical solutions of the EL equation.

Before presenting the numerical results, I briefly describe how the algorithms are implemented. To learn $L_d(\psi_i, \psi_{i+1})$, a neural network can be set up. Since it has only two inputs and one output, a deep network may not be necessary. For these two specific examples, the functional approximation for $L_d(\psi_i, \psi_{i+1})$ is implemented using polynomials in terms of $s \equiv \psi_i + \psi_{i+1}$ and $d \equiv \psi_{i+1} - \psi_i$, i.e.,

$$L_d(\psi_i, \psi_{i+1}) = \sum_{p=0}^P \sum_{q=0}^Q a_{pq} d^p s^q, \tag{12}$$

where a_{pq} are trainable parameters. For these two examples, I choose $(P, Q) = (4, 8)$, and the total number of trainable parameters are 45. For high-dimensional or vector discrete field theories, such as the Kepler problem in “Kepler problem” section, deep neural networks are probably more effective.

Example 1 The training data are plotted in Fig. 2 using empty square markers. It is a time sequence $\bar{\psi}_i$ ($i = 0, \dots, 50$) generated by the nonlinear ODE

$$2(\sin \psi + 1)\psi'' + (\psi')^2 \cos \psi + \frac{\pi^2}{200}\psi = 0 \tag{13}$$

with initial conditions $\psi(t = 0) = 1.2$ and $\psi'(t = 0) = 0$. Here ψ' denote $d\psi/dt$. The Lagrangian density for the system is

$$L(\psi, \psi') = (1 + \sin \psi)(\psi')^2 - \frac{\pi^2}{400}\psi^2. \tag{14}$$

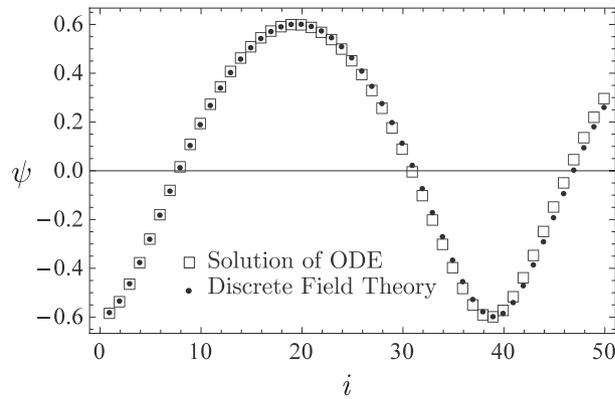


Figure 3. The predicted time sequence (solid circles) agrees with the time sequence (empty squares) accurately solved for from the nonlinear ODE (13). The dynamics starts at $\psi_0 = -0.6$, and its characteristics is significantly different from that of the time sequence in Fig. 2.

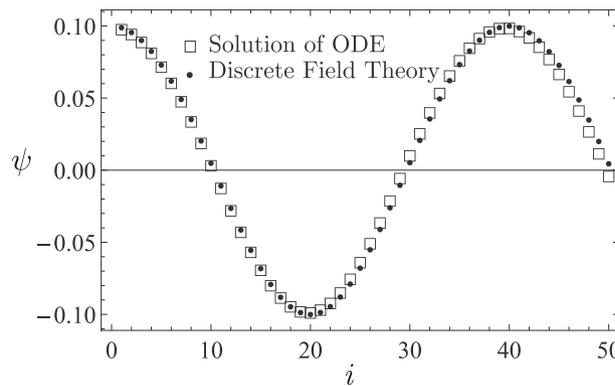


Figure 4. Starting at a much smaller amplitude, i.e., $\psi_0 = 0.1$, the predicted sequence (solid circles) shows the behavior of linear oscillation, in contrast with the strong nonlinearity of the sequence in Fig. 2 and the mild nonlinearity of the sequence in Fig. 3. The predicted time sequence agrees with the time sequence (empty squares) accurately solved for from the nonlinear ODE (13).

The optimizer for training the discrete Lagrangian density L_d is Newton’s algorithm with step-lengths reduced according to the amplitude of loss function. The discrete Lagrangian density L_d is trained until the loss function on $\bar{\psi}$ is less than 10^{-7} , then it is served. Plotted in Fig. 2 using solid circle markers are the predicted time sequence ψ_i using the initial conditions of the training data, i.e., $\psi_0 = \bar{\psi}_0$ and $\psi_1 = \bar{\psi}_1$. The predicted sequence ψ_i and the training sequence $\bar{\psi}_i$ are barely distinguishable in the figure.

The learned discrete field theory is then served with two sets of new initial conditions, and the predicted time sequences are plotted using solid circle markers in Figs. 3 and 4 against the time sequences solved for from the nonlinear ODE (13). The predicted sequence in Fig. 3 starts at $\psi_0 = -0.6$, and its dynamic characteristics is significantly different from that of the sequence in Fig. 2. The predicted sequence in Fig. 4 starts at a much smaller amplitude, i.e., $\psi_0 = 0.1$, and shows the behavior of linear oscillation, in contrast with the strong nonlinearity of the sequence in Fig. 2 and the mild nonlinearity of the sequence in Fig. 3. The agreement between the predictions of the learned discrete field theory and the accurate solutions of the nonlinear ODE (13) is satisfactory. These numerical results demonstrate that the proposed algorithms for machine learning and serving of discrete field theories are effective in terms of capturing the structure and predicting the dynamical behavior of the physical system.

Example 2 The training data are plotted in Fig. 5 using empty square markers. It is a time sequence $\bar{\psi}_i$ ($i = 0, \dots, 50$) generated by the nonlinear ODE

$$\psi'' - 0.03[\sin(1 - \psi^2)\psi + 0.1] = 0 \tag{15}$$

with initial conditions $\psi(t = 0) = 1.7$ and $\psi'(t = 0) = 0$. The Lagrangian for the system is

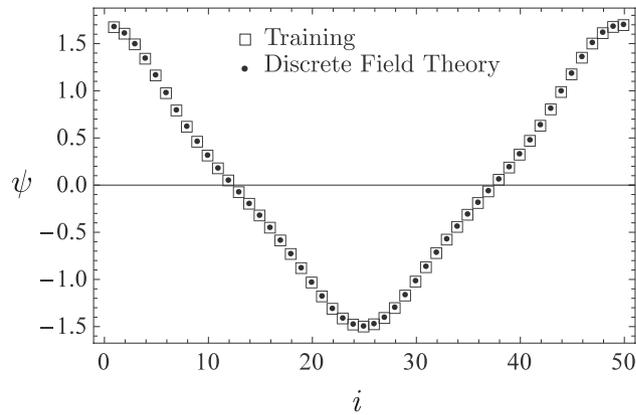


Figure 5. The training sequence (empty squares) represents a nonlinear oscillation in potential wall between $\psi = \pm 1.6$ in Fig. 6. The trained discrete Lagrangian density L_d is accepted when the loss function $F(\bar{\psi})$ on the training sequence is less than 10^{-7} . The predicted sequence (solid circles) from the learned discrete field theory agrees very well with the training sequence.

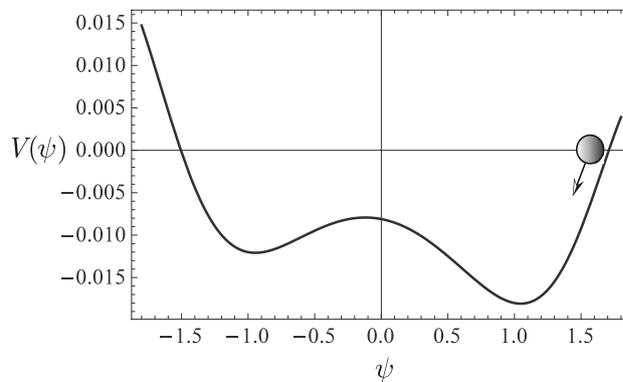


Figure 6. The training sequence in Fig. 5 represents a nonlinear oscillation in the large potential wall between $\psi = \pm 1.6$. There are two small potential wells secondary to the large potential well, one on the left between $\psi = -1.3$ and $\psi = -0.1$, and one on the right between $\psi = -0.1$ and $\psi = 1.5$.

$$L(\psi, \psi') = \frac{1}{2}(\psi')^2 - V(\psi), \tag{16}$$

$$V(\psi) = -0.015[\cos(\psi^2 - 1) + 0.2\psi], \tag{17}$$

where $V(\psi)$ is a nonlinear potential plotted in Fig. 6. The training sequence represents a nonlinear oscillation in the potential well between $\psi = \pm 1.6$. The trained discrete Lagrangian density L_d is accepted when the loss function $F(\bar{\psi})$ on the training sequence is less than 10^{-7} . The predicted sequence ψ_i (solid circles in Fig. 5) by the serving algorithm from the learned discrete field theory agrees very well with the training sequence $\bar{\psi}_i$.

The learned discrete field theory predicts two very different types of dynamical sequences shown in Figs. 7 and 8. The predicted sequences are plotted using solid circle markers and the sequences accurately solved for from the nonlinear ODE (15) are plotted using empty square markers. The sequence predicted in Fig. 7 is a nonlinear oscillation in the small potential well between $\psi = -0.1$ and $\psi = 1.5$ on the right of Fig. 6, and the sequence predicted in Fig. 8 is a nonlinear oscillation in the small potential well between $\psi = -1.3$ and $\psi = -0.1$ on the left. For both cases, the predictions of the learned discrete field theory agree with the accurate solutions. Observe that in Fig. 6 the two small potential wells are secondary to the large potential wall between $\psi = \pm 1.6$. In Fig. 5 the small-scale fluctuations in the training sequence, which is a nonlinear oscillation in the large potential well, encode the structures of the small potential wells. The training algorithm is able to diagnose and record these fine structures in the learned discrete Lagrangian density, and the serving algorithm correctly predicts the secondary dynamics due to them.

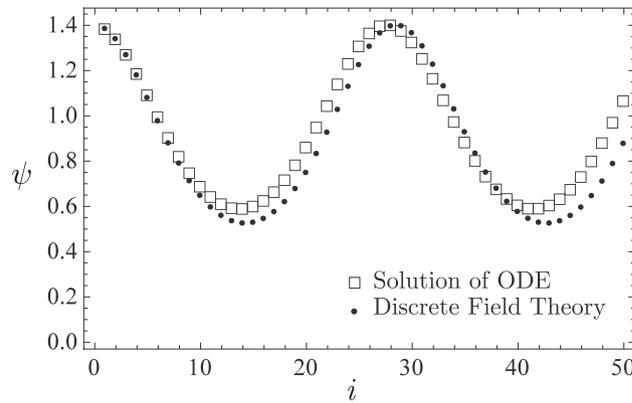


Figure 7. The learned discrete field theory correctly predicts a nonlinear oscillation in the small potential well between $\psi = -0.1$ and $\psi = 1.5$ on the right of Fig. 6. The predicted sequence (solid circles) agrees with the accurate solution (empty squares) of the nonlinear ODE (15).

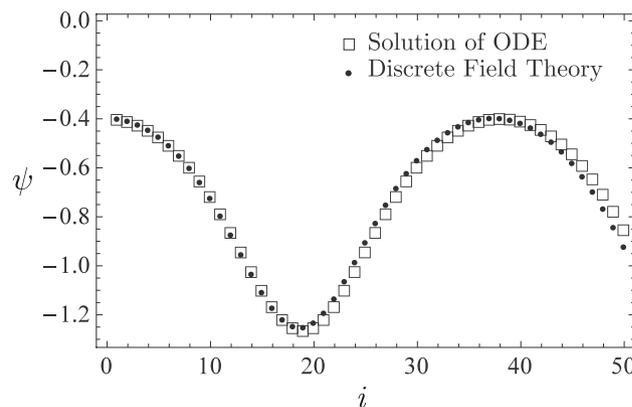


Figure 8. The learned discrete field theory correctly predicts a nonlinear oscillation in the small potential well between $\psi = -1.3$ and $\psi = -0.1$ on the left of Fig. 6. The predicted sequence (solid circles) agrees with the accurate solution (empty squares) of the nonlinear ODE (15).

Kepler problem

In this section, to further demonstrate the effectiveness of the method developed, I apply it to the Kepler problem, which is concerned with dynamics of planetary orbits in the solar system. Let’s turn the clock back to 1601, when Kepler inherited the observational data of planetary orbits meticulously collected by his mentor Tycho Brahe. It took Kepler 5 years to discover his first and second laws of planetary motion, and another 78 years before Newton solved the Kepler problem using his laws of motion and universal gravitation⁸⁴. Assume that we have a set of data similar to that of Kepler, as displayed in Fig. 9. For simplicity, the data are the orbits of the Mercury, Venus, Earth, Mars, Ceres and Jupiter generated by solving Newton’s equation of motion for a planet in the gravity field of the Sun according to Newton’s law of universal gravitation. The spatial and temporal normalization scale-lengths are 1 a.u. and 58.14 days, respectively, and the time-steps of the orbital data is 0.05.

My goal here is not to rediscover Kepler’s laws of planetary motion or Newton’s laws of motion and universal gravitation by machine learning. Instead, I train a discrete field theory from the orbits displayed in Fig. 9 and then serve it to predict new planetary orbits. For this case, the discrete field theory is about a 2D vector field defined on the time grid. Denote the field as $\psi_i = (x_i, y_i)$, where i is the index for the time grid, and x_i and y_i are the 2D coordinates of a planet in the solar system. In terms of the discrete field, the discrete Lagrangian density L_d is a function on \mathbb{R}^4 ,

$$L_d(\psi_i, \psi_{i+1}) = L_d(x_i, y_i, x_{i+1}, y_{i+1}), \tag{18}$$

and the DEL is a vector equation with two components,

$$EL_{x_i} = \frac{\partial L_d(x_{i-1}, y_{i-1}, x_i, y_i)}{\partial x_i} + \frac{\partial L_d(x_i, y_i, x_{i+1}, y_{i+1})}{\partial x_i} = 0, \tag{19}$$

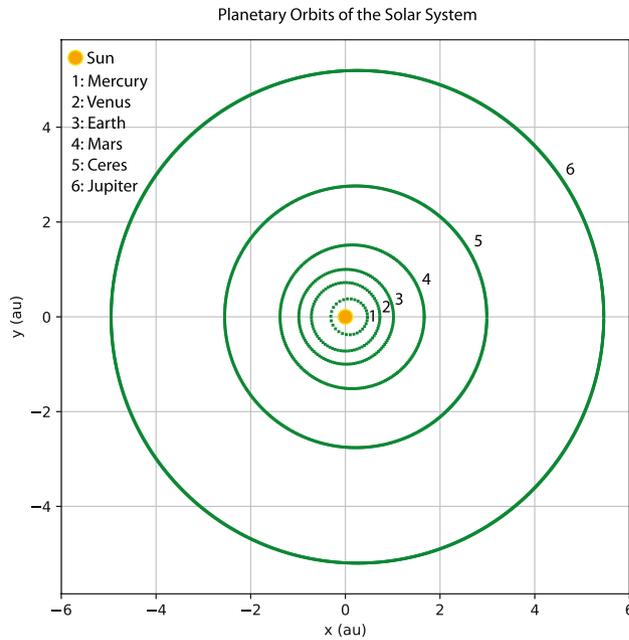


Figure 9. Orbits of the Mercury, Venus, Earth, Mars, Ceres and Jupiter generated by solving Newton’s equation of motion for a planet in the gravity field of the Sun according to Newton’s law of universal gravitation. These orbits are the training data for the discrete field theory.

$$EL_{y_i} = \frac{\partial L_d(x_{i-1}, y_{i-1}, x_i, y_i)}{\partial y_i} + \frac{\partial L_d(x_i, y_i, x_{i+1}, y_{i+1})}{\partial y_i} = 0. \tag{20}$$

The loss function on a data set $\psi = (x, y)$ is

$$F(x, y) = \frac{1}{I} \sum_{i=1}^{I-1} [EL_{x_i}(x, y)^2 + EL_{y_i}(x, y)^2].$$

Akin to the situation in “Examples of learning and predicting nonlinear oscillations” section, the serving algorithm preserves exactly an discrete symplectic form defined by

$$\Omega(x_i, y_i, x_{i+1}, y_{i+1}) = d\theta, \tag{21}$$

$$\theta = \frac{\partial L_d(x_i, y_i, x_{i+1}, y_{i+1})}{\partial x_{i+1}} dx_{i+1} + \frac{\partial L_d(x_i, y_i, x_{i+1}, y_{i+1})}{\partial y_{i+1}} dy_{i+1}. \tag{22}$$

To model the discrete Lagrangian density $L_d(x_i, y_i, x_{i+1}, y_{i+1})$, I use a fully connected neural network with two hidden layers, each of which has 40 neurons with the sigmoid activation function. The network is randomly initialized with a normal distribution, and then trained by the Adam optimizer⁸² until the averaged loss on a single time grid-point is reduced by a factor of 10^5 relative to its initial value. Starting from the same initial conditions as the training orbits, the serving algorithm of the trained discrete field theory predicts the orbits plotted using red markers in Fig. 10 against the training orbits indicated by green markers. The agreement between the predicted and training orbits shown in the figure validates the discrete field theory learned. To serve it for the purpose of predicting new orbits, let’s consider the scenario of launching a device at the Perihelion of the Earth orbit with an orbital velocity v_p larger than that of the Earth. Four such orbits, labeled by e1, e2, h, and p with $v_p = 1.13, 1.26, 1.40,$ and 1.50 , are plotted in Fig. 11 along with the orbit of the Earth, which is the inner most ellipse labeled by e0 with $v_p = 0.98$. Orbits plotted using red markers are predictions of the trained discrete field theory, and orbits plotted using blue markers are solutions according to Newton’s laws of motion and gravitation. The agreement is satisfactory. Orbits e1 and e2 are elliptical, and Orbit p is the parabolic escaping orbit and Orbit h is the hyperbolic escaping orbit.

Similar study is carried out for the orbits initiated from the Perihelion of the Mercury orbit with an orbital velocities v_p larger than that of the Mercury. Four such orbits are shown in Fig. 12 using the same plotting markers and labels as in Fig. 11. The inner most elliptical orbit is that of the Mercury with $v_p = 1.30$. The orbit velocities at the Perihelion of the other four orbits are $v_p = 1.56, 1.80, 2.07, 2.20$. Again, the the predictions of the trained discrete field theory agree well with those of Newton’s laws.

It is noteworthy that the trained discrete field theory correctly predicts the parabolic and hyperbolic escaping orbits, even though the training orbits are all elliptical, see Figs. 9 and 10. Historically, Kepler argued that escaping orbits and elliptical orbits are governed by different laws. It was Newton who discovered or “learned”

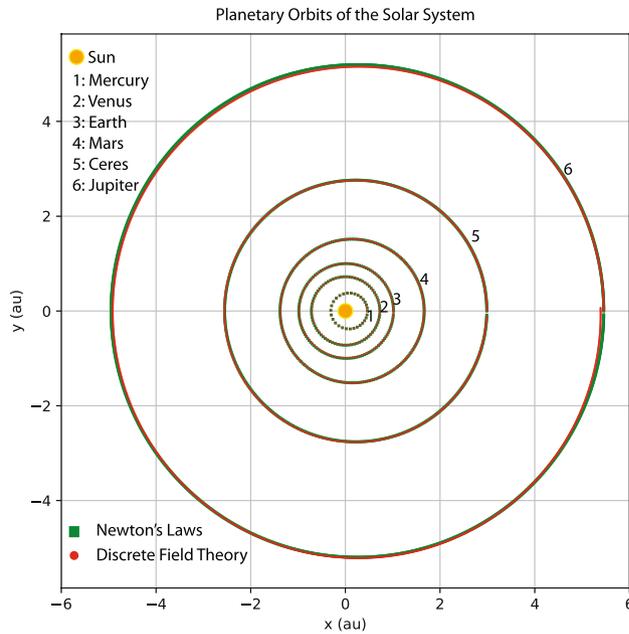


Figure 10. Orbits of the Mercury, Venus, Earth, Mars, Ceres and Jupiter. The orbits indicated by red markers are generated by the learned discrete field theory. The orbits indicated by green markers are the training orbits from Fig. 9.

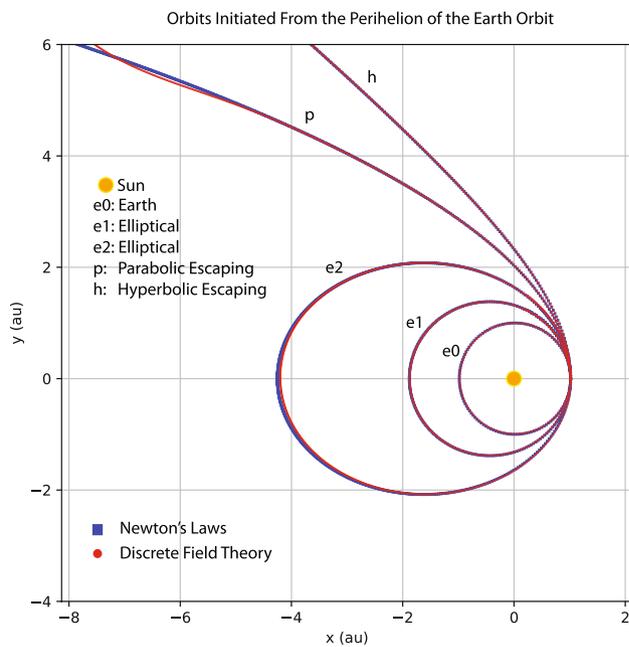


Figure 11. Orbits initiated from the Perihelion of the Earth orbit with initial velocities $v_p = 0.98, 1.13, 1.26, 1.40,$ and 1.50 . Orbit e_0 is the Earth orbit. Orbits e_1 and e_2 are elliptical, and Orbit p is the parabolic escaping orbit and Orbit h is the hyperbolic escaping orbit. Red markers are the predictions of the trained discrete field theory, and blue markers are solutions according to Newton's laws of motion and universal gravitation.

the $1/r$ dependency of the gravitational field from Kepler's laws of planetary motion and Tycho Brahe's data, and unified the elliptical orbits and escaping orbits under the same law of physics. Newton's methodology has a deep and long-lasting influence on physicists after him. The results displayed in Figs. 11 and 12 show that the machine learning and serving algorithms solve the Kepler problem in terms of correctly prediction planetary orbits without knowing or learning Newton's laws of motion and universal gravitation.

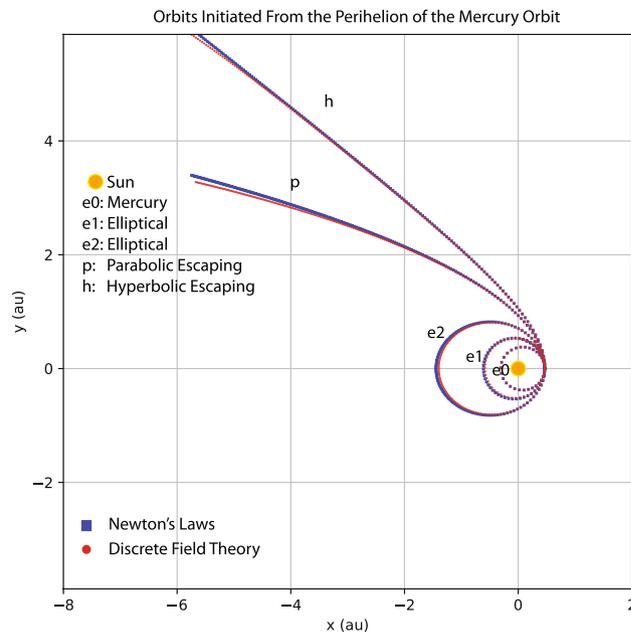


Figure 12. Orbits initiated from the Perihelion of the Mercury orbit with initial velocities $v_p = 1.30, 1.56, 1.80, 2.07,$ and 2.20 . Plotting markers and labels are similar to those in Fig. 11.

The algorithms developed are robust against variations of the governing laws of physics, because the method does not require any knowledge of the laws of physics other than the fundamental assumption that the governing laws are field theories. When the effects of special relativity or general relativity are important, the algorithms are expected to be valid as well.

Nevertheless, a few footnotes are in order. (i) There exist small discrepancies between the predictions from the learned discrete field theory and Newton's laws in Figs. 11 and 12 when $r = \sqrt{x^2 + y^2} \gtrsim 7$. This is because no training orbit in this domain is provided to the learning algorithm. The orbits predicted there are thus less accurate. To further understand the effect of reduced training data, I remove the orbit of the Ceres from the training data set and then apply the trained discrete Lagrangian to predict orbits initiated from the Perihelion of the Earth orbit as in Fig. 11. The result is displayed in Fig. 13, which shows a larger deviation from the prediction of Newton's laws compared with Fig. 11. The increased discrepancies are understandable because, without the orbital data of the Ceres, there is no training data covering the space of $1.8 \lesssim r \lesssim 4.5$. This indicates that trained discrete Lagrangian cannot accurately interpolate the gravitational field to regions without training data. On the other hand, without any additional physical assumption, there is no unique way to determine the gravitational field in a region without observational data. For example, this region could be vacuum or other small celestial bodies may exist and modify the gravitational field there. The neural network model of the discrete Lagrangian has the flexibility to learn the gravitational field in the region, whatever it might be, once training data in the region is available. But without the training data, the gravitational field in the region is under-determined in the neural network model. The calculation based on Newton's laws implicitly assumes that the region is vacuum. Hence the larger discrepancies. (ii) The study presented above is meant to be a proof of principle. Practical factors, such as the noise of the training data and the three-body effects, are not included. Here, I would like to present a preliminary study on the effect of noise. In the orbital training data displayed in Fig. 9, I now add in a Gaussian noise with zero mean and 10^{-3} standard deviation. To denoise, the algorithm first feeds the data through a simple low-pass filter before training the discrete Lagrangian. The parameters of the filter are varied to minimize the loss function. The predictions of the trained discrete Lagrangian for the orbits initiated from the Perihelion of the Earth orbit are shown in Fig. 14. Comparing Figs. 11 and 14, we observe that the Gaussian noise at this level does not alter the predicted orbits substantially. However, the resulting error is not negligible and it is desirable to improve the robustness of the machine learning model against noise. We plan to carry out a systematic study on this topic. Several techniques are being considered. For instance, a noise-canceling signal can be searched to minimize the loss function during the training of the discrete Lagrangian. Generative machine learning models, which have found many applications in physics recently^{23,41,43,85,86}, can be applied for the purpose of denoising as well^{85,87}.

Conclusions and discussion

In this paper, a method for machine learning and serving of discrete field theories in physics is developed. The learning algorithm trains a discrete field theory from a set of observational data of the field on a spacetime lattice, and the serving algorithm employs the learned discrete field theory to predict new observations of the field for given new boundary and initial conditions.

The algorithm does not attempt to capture statistical properties of the training data, nor does it try to discover differential equations that govern the training data. Instead, it learns a discrete field theory that underpins

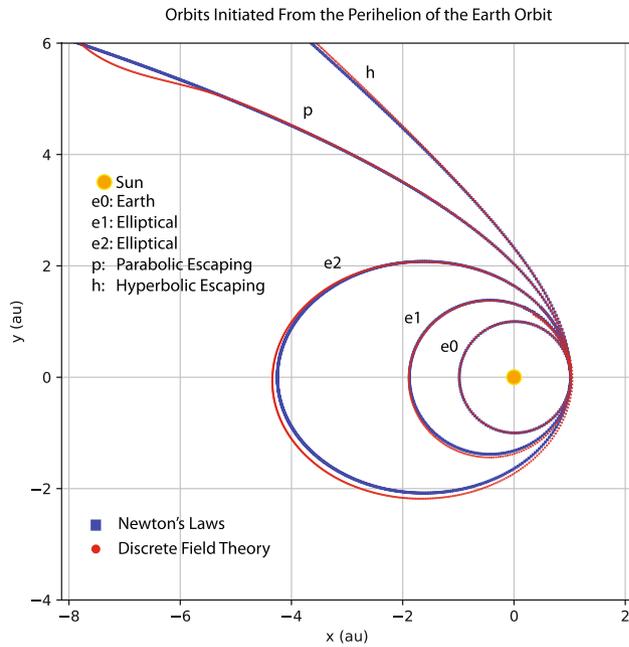


Figure 13. The effect of reduced training data. The orbit of the Ceres is removed from the training data shown in Fig. 9. The simulated orbits are the same as in Fig. 11. Red markers are the predictions of the trained discrete field theory, and blue markers are solutions according to Newton's laws of motion and universal gravitation.

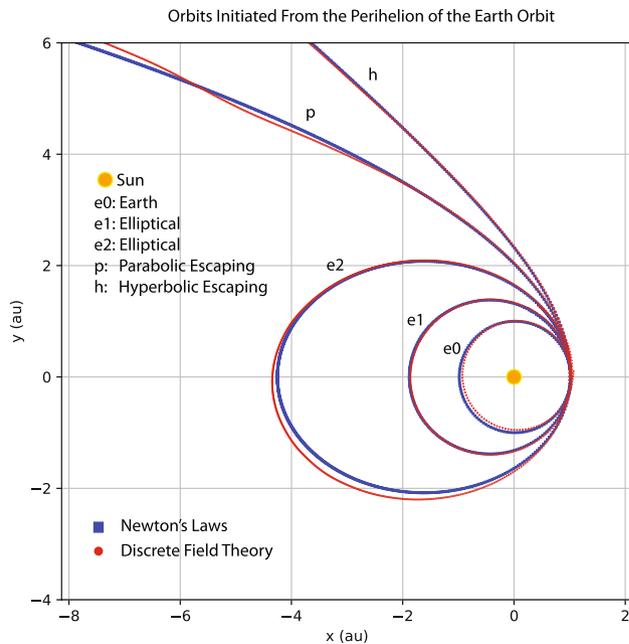


Figure 14. The effect of noise in training data. A Gaussian noise with zero mean and 10^{-3} standard deviation is added to the training data shown in Fig. 9. The simulated orbits are the same as in Fig. 11. Red markers are the predictions of the trained discrete field theory, and blue markers are solutions according to Newton's laws of motion and universal gravitation.

the observed field. Because the learned field theory is discrete, it overcomes the difficulties associated with the learning of continuous theories. Compared with continuous field theories, discrete field theories can be served more easily and with improved long-term accuracy and fidelity. The serving algorithm of discrete field theories belongs to the family of structure-preserving geometric algorithms⁴⁶⁻⁷⁵, which have been proven to be superior to the conventional algorithms based on discretization of differential equations. The demonstrated advantages of

discrete field theories relative to continuous theories in terms of machine learning compatibility are consistent with Bostrom's simulation hypothesis⁷⁸.

I should point out that in general, a specific machine learning algorithm is often more effective for certain types of data. The data relevant to the present study are assumed to be observations of physical fields in spacetime governed by field theories, even though laws of physics in specific forms, such as Newton's laws of motion and gravity, are not needed for the algorithms developed in the present study to be effective in terms of correctly predicting observations. Admittedly, the assumption of the existence of an underpinning field theory puts a strong constraint on the applicability of the method. Many physics and engineering problems do not have field theoretical formulations. Most non-conservative systems fall into this category. The algorithms developed in the present study certainly do not apply in these situations. Further investigation is needed to incorporate non-conservative effects.

In addition, the machine learning community has developed large amounts of innovative methods and techniques, many of which can be explored to facilitate the machine learning of discrete field theories. For instance, if a data set is generated by systems governed by different Lagrangians, we can include both the observations and the corresponding discrete Lagrangians as the training data for a generative model^{23,41,43,85,86} using the method of variational autoencoder⁸⁸. After training, the encoder establishes a map between observations and discrete Lagrangians. And the space of discrete Lagrangian can be continuously sampled to generate new Lagrangians and predict new observations. The generative model can also be utilized to remove noise in the training data, as discussed in "Kepler problem" section.

We plan to pursue these topics in the follow-up study.

Received: 29 July 2020; Accepted: 27 October 2020

Published online: 09 November 2020

References

- Narendra, K. S. & Parthasarathy, K. Identification and control of dynamical systems using neural networks. *IEEE Trans. Neural Netw.* **1**, 4–27 (1990).
- Narendra, K. S. & Parthasarathy, K. Neural networks and dynamical systems. *Int. J. Approx. Reason.* **6**, 109–131 (1992).
- Ramacher, U. Hamiltonian dynamics of neural networks. *Neural Netw.* **6**, 547–557 (1993).
- Howse, J. W., Abdallah, C. T. & Heileman, G. L. Gradient and Hamiltonian dynamics applied to learning in neural networks. *Adv. Neural Inf. Process. Syst.* **8**, 274–280 (1995).
- Wilde, P. D. Class of Hamiltonian neural networks. *Phys. Rev. E* **47**, 1392–1396 (1993).
- Weinan, E. A proposal on machine learning via dynamical systems. *Commun. Math. Stat.* **5**, 1–11 (2017).
- Chen, T. Q., Rubanova, Y., Bettencourt, J. & Duvenaud, D. Neural ordinary differential equations. *Adv. Neural Inf. Process. Syst.* **31**, 6572–6583 (2018).
- Haber, E. & Ruthotto, L. Stable architectures for deep neural networks. *Inverse Probl.* **34**, 014004 (2018).
- Wroblewski, D., Jahns, G. & Leuer, J. Tokamak disruption alarm based on a neural network model of the high-beta limit. *Nucl. Fusion* **37**, 725–741 (1997).
- Vannucci, A., Oliveira, K. & Tajima, T. Forecast of TEXT plasma disruptions using soft x rays as input signal in a neural network. *Nucl. Fusion* **39**, 255–262 (1999).
- Yoshino, R. Neural-net disruption predictor in JT-60u. *Nucl. Fusion* **43**, 1771–1786 (2003).
- Kates-Harbeck, J., Svyatkovskiy, A. & Tang, W. Predicting disruptive instabilities in controlled fusion plasmas through deep learning. *Nature* **568**, 526 (2019).
- Wu, J.-L., Xiao, H. & Paterson, E. Physics-informed machine learning approach for augmenting turbulence models: A comprehensive framework. *Phys. Rev. Fluids* **3**, 074602 (2018).
- Bar-Sinai, Y., Hoyer, S., Hickey, J. & Brenner, M. P. Learning data-driven discretizations for partial differential equations. *Proc. Natl. Acad. Sci.* **116**, 15344–15349 (2019).
- Han, J., Ma, C., Ma, Z. & Weinan, E. Uniformly accurate machine learning-based hydrodynamic models for kinetic equations. *Proc. Natl. Acad. Sci.* **116**, 21983–21991 (2019).
- Dissanayake, M. W. M. G. & Phan-Thien, N. Neural-network-based approximations for solving partial differential equations. *Commun. Numer. Methods Eng.* **10**, 195–201 (1994).
- Meade, A. J. Jr. & Fernández, A. A. The numerical solution of linear ordinary differential equations by feedforward neural networks. *Math. Comput. Model.* **19**, 1–25 (1994).
- Meade, A. J. Jr. & Fernández, A. A. Solution of nonlinear ordinary differential equations by feedforward neural networks. *Math. Comput. Model.* **20**, 19–44 (1994).
- Lagaris, I. E., Likas, A. & Fotiadis, D. I. Artificial neural networks for solving ordinary and partial differential equations. *IEEE Trans. Neural Netw.* **9**, 987–1000 (1998).
- Bailer-Jones, C. A. L., MacKay, D. J. C. & Withers, P. J. A recurrent neural network for modelling dynamical systems. *Netw. Comput. Neural Syst.* **9**, 531–547 (1998).
- Long, Z., Lu, Y. & Dong, B. PDE-net 2.0: Learning PDEs from data with a numeric-symbolic hybrid deep network. *J. Comput. Phys.* **399**, 108925 (2019).
- Carleo, G. & Troyer, M. Solving the quantum many-body problem with artificial neural networks. *Science* **355**, 602–606 (2017).
- Nomura, Y., Darmawan, A. S., Yamaji, Y. & Imada, M. Restricted boltzmann machine learning for solving strongly correlated quantum systems. *Phys. Rev. B* **96**, 205152 (2017).
- Bongard, J. & Lipson, H. Automated reverse engineering of nonlinear dynamical systems. *Proc. Natl. Acad. Sci.* **104**, 9943–9948 (2007).
- Schmidt, M. & Lipson, H. Distilling free-form natural laws from experimental data. *Science* **324**, 81–85 (2009).
- Brunton, S. L., Proctor, J. L. & Kutz, J. N. Discovering governing equations from data by sparse identification of nonlinear dynamical systems. *Proc. Natl. Acad. Sci.* **113**, 3932–3937 (2016).
- Rudy, S. H., Brunton, S. L., Proctor, J. L. & Kutz, J. N. Data-driven discovery of partial differential equations. *Sci. Adv.* **3**, e1602614 (2017).
- Schaeffer, H. Learning partial differential equations via data discovery and sparse optimization. *Proc. R. Soc. A Math. Phys. Eng. Sci.* **473**, 20160446 (2017).
- Baydin, A. G., Pearlmutter, B. A., Radul, A. A. & Siskind, J. M. Automatic differentiation in machine learning: A survey. *J. Mach. Learn. Res.* **18**, 153:1–153:43 (2017).

30. Raissi, M. Deep hidden physics models: Deep learning of nonlinear partial differential equations. *J. Mach. Learn. Res.* **19**, 1–24 (2018).
31. Cranmer, M. D., Xu, R., Battaglia, P. & Ho, S. Learning symbolic physics with graph networks. Preprint at [arXiv:1909.05862v2](https://arxiv.org/abs/1909.05862v2) (2019).
32. Gellß, P., Klus, S., Eisert, J. & Schßtte, C. Multidimensional approximation of nonlinear dynamical systems. *J. Comput. Nonlinear Dyn.* **14**, 061006 (2019).
33. Raissi, M., Perdikaris, P. & Karniadakis, G. Physics-informed neural networks: A deep learning framework for solving forward and inverse problems involving nonlinear partial differential equations. *J. Comput. Phys.* **378**, 686–707 (2019).
34. Wu, K., Qin, T. & Xiu, D. Structure-preserving method for reconstructing unknown Hamiltonian systems from trajectory data. Preprint at [arXiv:1905.10396v1](https://arxiv.org/abs/1905.10396v1) (2019).
35. Lutter, M., Ritter, C. & Peters, J. Deep Lagrangian networks: Using physics as model prior for deep learning. Preprint at [arXiv:1907.04490v1](https://arxiv.org/abs/1907.04490v1) (2019).
36. Bertalan, T., Dietrich, F., Meziç, I. & Kevrekidis, I. G. On learning Hamiltonian systems from data. Preprint at [arXiv:1907.12715v2](https://arxiv.org/abs/1907.12715v2) (2019).
37. Greydanus, S., Dzamba, M. & Yosinski, J. Hamiltonian neural networks. Preprint at [arXiv:1906.01563v3](https://arxiv.org/abs/1906.01563v3) (2019).
38. Zhong, Y. D., Dey, B. & Chakraborty, A. Symplectic ODE-net: Learning Hamiltonian dynamics with control. Preprint at [arXiv:1909.12077v1](https://arxiv.org/abs/1909.12077v1) (2019).
39. Sanchez-Gonzalez, A., Bapst, V., Cranmer, K. & Battaglia, P. Hamiltonian graph networks with ODE integrators. Preprint at [arXiv:1909.12790v1](https://arxiv.org/abs/1909.12790v1) (2019).
40. Chen, Z., Zhang, J., Arjovsky, M. & Bottou, L. Symplectic recurrent neural networks. Preprint at [arXiv:1909.13334v1](https://arxiv.org/abs/1909.13334v1) (2019).
41. Toth, P. *et al.* Hamiltonian generative networks. Preprint at [arXiv:1909.13789v1](https://arxiv.org/abs/1909.13789v1) (2019).
42. Shanahan, P. E., Trewartha, D. & Detmold, W. Machine learning action parameters in lattice quantum chromodynamics. *Phys. Rev. D* **97**, 094506 (2018).
43. Noé, F., Olsson, S., Köhler, J. & Wu, H. Boltzmann generators: Sampling equilibrium states of many-body systems with deep learning. *Science* **365**, eaaw1147 (2019).
44. Halverson, J., Nelson, B. & Ruehle, F. Branes with brains: Exploring string vacua with deep reinforcement learning. *J. High Energy Phys.* **2019**, 3 (2019).
45. Cranmer, K., Golkar, S. & Pappadopulo, D. Inferring the quantum density matrix with machine learning. Preprint at [arXiv:1904.05903v1](https://arxiv.org/abs/1904.05903v1) (2019).
46. Feng, K. On difference schemes and symplectic geometry. In *The Proceedings of 1984 Beijing Symposium on Differential Geometry and Differential Equations* (ed. Feng, K.), 42–58 (Science Press, 1985).
47. Sanz-Serna, J. M. & Calvo, M. P. *Numerical Hamiltonian Problems* (Chapman and Hall, London, 1994).
48. Marsden, J. E., Patrick, G. W. & Shkoller, S. Multisymplectic geometry, variational integrators, and nonlinear pdes. *Commun. Math. Phys.* **199**, 351–395 (1998).
49. Marsden, J. E. & West, M. Discrete mechanics and variational integrators. *Acta Numerica* **10**, 357–514 (2001).
50. Hairer, E., Lubich, C. & Wanner, G. *Geometric Numerical Integration: Structure-preserving Algorithms for Ordinary Differential Equations* Vol. 31 (Springer, Berlin, 2006).
51. Qin, H. & Guan, X. Variational symplectic integrator for long-time simulations of the guiding-center motion of charged particles in general magnetic fields. *Phys. Rev. Lett.* **100**, 035006 (2008).
52. Squire, J., Qin, H. & Tang, W. M. Geometric integration of the Vlasov-Maxwell system with a variational particle-in-cell scheme. *Phys. Plasmas* **19**, 084501 (2012).
53. Xiao, J., Liu, J., Qin, H. & Yu, Z. A variational multi-symplectic particle-in-cell algorithm with smoothing functions for the Vlasov-Maxwell system. *Phys. Plasmas* **20**, 102517 (2013).
54. Zhang, R. *et al.* Canonicalization and symplectic simulation of the gyrocenter dynamics in time-independent magnetic fields. *Phys. Plasmas* **21**, 032504 (2014).
55. Zhou, Y., Qin, H., Burby, J. W. & Bhattacharjee, A. Variational integration for ideal magnetohydrodynamics with built-in advection equations. *Phys. Plasmas* **21**, 102109 (2014).
56. He, Y. *et al.* Hamiltonian time integrators for Vlasov-Maxwell equations. *Phys. Plasmas* **22**, 124503 (2015).
57. Xiao, J., Liu, J., Qin, H., Yu, Z. & Xiang, N. Variational symplectic particle-in-cell simulation of nonlinear mode conversion from extraordinary waves to Bernstein waves. *Phys. Plasmas* **22**, 092305 (2015).
58. Ellison, C. L., Finn, J. M., Qin, H. & Tang, W. M. Development of variational guiding center algorithms for parallel calculations in experimental magnetic equilibria. *Plasma Phys. Control. Fusion* **57**, 054007 (2015).
59. Qin, H. *et al.* Canonical symplectic particle-in-cell method for long-term large-scale simulations of the Vlasov-Maxwell equations. *Nucl. Fusion* **56**, 014001 (2016).
60. He, Y., Sun, Y., Qin, H. & Liu, J. Hamiltonian particle-in-cell methods for Vlasov-Maxwell equations. *Phys. Plasmas* **23**, 092108 (2016).
61. Xiao, J. *et al.* Explicit high-order noncanonical symplectic algorithms for ideal two-fluid systems. *Phys. Plasmas* **23**, 112107 (2016).
62. Zhang, R. *et al.* Explicit symplectic algorithms based on generating functions for charged particle dynamics. *Phys. Rev. E* **94**, 013205 (2016).
63. Wang, Y., Liu, J. & Qin, H. Lorentz covariant canonical symplectic algorithms for dynamics of charged particles. *Phys. Plasmas* **23**, 122513 (2016).
64. Xiao, J., Qin, H., Liu, J. & Zhang, R. Local energy conservation law for a spatially-discretized Hamiltonian Vlasov-Maxwell system. *Phys. Plasmas* **24**, 062112 (2017).
65. Burby, J. W. Finite-dimensional collisionless kinetic theory. *Phys. Plasmas* **24**, 032101 (2017).
66. Chen, Q. *et al.* Canonical symplectic structure and structure-preserving geometric algorithms for Schrödinger-Maxwell systems. *J. Comput. Phys.* **349**, 441–452 (2017).
67. He, Y., Zhou, Z., Sun, Y., Liu, J. & Qin, H. Explicit k-symplectic algorithms for charged particle dynamics. *Phys. Lett. A* **381**, 568–573 (2017).
68. Burby, J. W. & Ellison, C. L. Toroidal regularization of the guiding center Lagrangian. *Phys. Plasmas* **24**, 110703 (2017).
69. Kraus, M., Kormann, K., Morrison, P. J. & Sonnendrücker, E. GEMPIC: geometric electromagnetic particle-in-cell methods. *J. Plasma Phys.* **83**, 905830401 (2017).
70. Xiao, J., Qin, H. & Liu, J. Structure-preserving geometric particle-in-cell methods for Vlasov-Maxwell systems. *Plasma Sci. Technol.* **20**, 110501 (2018).
71. Ellison, C. L. *et al.* Degenerate variational integrators for magnetic field line flow and guiding center trajectories. *Phys. Plasmas* **25**, 052502 (2018).
72. Xiao, J., Qin, H., Shi, Y., Liu, J. & Zhang, R. A lattice Maxwell system with discrete space-time symmetry and local energy-momentum conservation. *Phys. Lett. A* **383**, 808–812 (2019).
73. Xiao, J. & Qin, H. Field theory and a structure-preserving geometric particle-in-cell algorithm for drift wave instability and turbulence. *Nucl. Fusion* **59**, 106044 (2019).
74. Xiao, J. & Qin, H. Explicit high-order gauge-independent symplectic algorithms for relativistic charged particle dynamics. *Comput. Phys. Commun.* **241**, 19–27 (2019).

75. Glasser, A. S. & Qin, H. The geometric theory of charge conservation in particle-in-cell simulations. *J. Plasma Phys.* **86**, 835860303 (2020).
76. de Maupertuis, P. Accord de différentes lois de la nature qui avaient jusqu'ici paru incompatibles. *Mém. As. Sc. Paris* 417 (1744).
77. de Maupertuis, P. Le lois de mouvement et du repos, déduites d'un principe de métaphysique. *Mém. Ac. Berlin* 267 (1746).
78. Bostrom, N. Are we living in a computer simulation?. *Philos. Quart.* **53**, 243–255 (2003).
79. Beane, S. R., Davoudi, Z. & Savage, M. J. Constraints on the universe as a numerical simulation. *Eur. Phys. J. A* **50**, 148 (2014).
80. Glasser, A. S. & Qin, H. Lifting spacetime's Poincaré symmetries. Preprint at [arXiv:1902.04395v1](https://arxiv.org/abs/1902.04395v1) (2019).
81. Glasser, A. S. & Qin, H. Restoring Poincaré symmetry to the lattice. Preprint at [arXiv:1902.04396v1](https://arxiv.org/abs/1902.04396v1) (2019).
82. Kingma, D. P. & Ba, J. Adam: A method for stochastic optimization. Preprint at [arXiv:1412.6980v9](https://arxiv.org/abs/1412.6980v9) (2014).
83. Davoudi, Z. & Savage, M. J. Restoration of rotational symmetry in the continuum limit of lattice field theories. *Phys. Rev. D* **86**, 054505 (2012).
84. Newton, I. *The Mathematical Papers of Isaac Newton*, Volume IV, 1684–1691 (Cambridge University Press, Cambridge, 2008).
85. Schawinski, K., Zhang, C., Zhang, H., Fowler, L. & Santhanam, G. K. Generative adversarial networks recover features in astrophysical images of galaxies beyond the deconvolution limit. *Monthly Notices of the Royal Astronomical Society: Letters* slx008 (2017).
86. Cerri, O., Nguyen, T. Q., Pierini, M., Spiropulu, M. & Vlimant, J.-R. Variational autoencoders for new physics mining at the large hadron collider. *J. High Energy Phys.* **2019**, 36 (2019).
87. Yang, Q. *et al.* Low-dose CT image denoising using a generative adversarial network with Wasserstein distance and perceptual loss. *IEEE Trans. Med. Imaging* **37**, 1348–1357 (2018).
88. Doersch, C. Tutorial on variational autoencoders. Preprint at [arXiv:1606.05908v2](https://arxiv.org/abs/1606.05908v2) (2016).

Acknowledgements

This research was supported by the U.S. Department of Energy (DE-AC02-09CH11466). I thank Alexander S. Glasser, Yichen Fu, Michael Churchill, George Wilkie, and Nick McGreivy for fruitful discussions.

Author contributions

H.Q. as a single author carried out all the work.

Competing interests

The author declares no competing interests.

Additional information

Correspondence and requests for materials should be addressed to H.Q.

Reprints and permissions information is available at www.nature.com/reprints.

Publisher's note Springer Nature remains neutral with regard to jurisdictional claims in published maps and institutional affiliations.



Open Access This article is licensed under a Creative Commons Attribution 4.0 International License, which permits use, sharing, adaptation, distribution and reproduction in any medium or format, as long as you give appropriate credit to the original author(s) and the source, provide a link to the Creative Commons licence, and indicate if changes were made. The images or other third party material in this article are included in the article's Creative Commons licence, unless indicated otherwise in a credit line to the material. If material is not included in the article's Creative Commons licence and your intended use is not permitted by statutory regulation or exceeds the permitted use, you will need to obtain permission directly from the copyright holder. To view a copy of this licence, visit <http://creativecommons.org/licenses/by/4.0/>.

This is a U.S. Government work and not under copyright protection in the US; foreign copyright protection may apply 2020