

## Author Correction: Testing sub-gravitational forces on atoms from a miniature in-vacuum source mass

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We miscalculated the gravitational acceleration arising from the tungsten mass by a factor of about two. The correct value is  $a_{\text{grav}} = (33 \pm 3) \text{ nm/s}^2$ . The corrected anomalous acceleration is  $\alpha_{\text{anomaly}} = a_{\text{grav}} - a_{\text{grav}} = (41 \pm 24) \text{ nm/s}^2$ . These values replace the erroneous ones of  $a_{\text{grav}} = (65 \pm 5) \text{ nm/s}^2$  and  $\alpha_{\text{anomaly}} = (9 \pm 24) \text{ nm/s}^2$ .

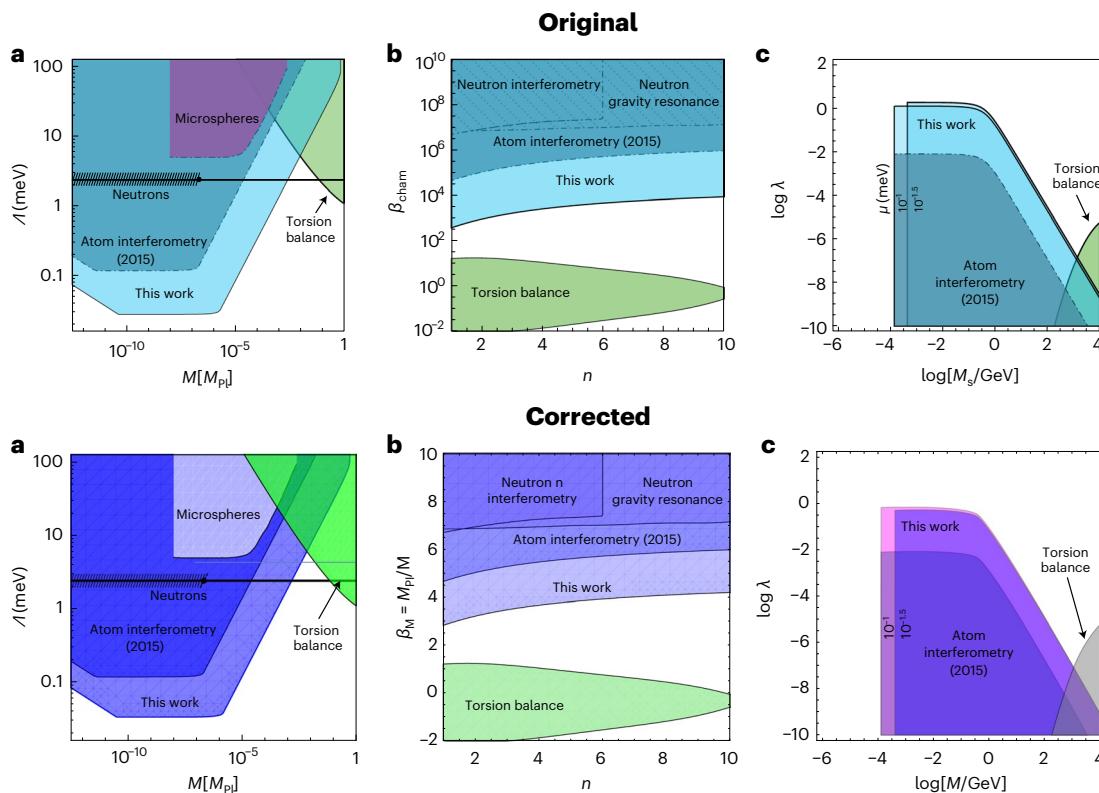
The corrected 95% confidence interval is  $-7 \text{ nm/s}^2 < \alpha_{\text{anomaly}} < 89 \text{ nm/s}^2$ . Using a one-tailed test to bound fifth-force interactions (which must be attractive for scalar fields with a universal matter coupling), we constrain anomalous accelerations  $\alpha_{\text{anomaly}} < 81 \text{ nm/s}^2$  (95% confidence level).

For chameleon fields with  $\Lambda = \Lambda_0 = 2.4 \text{ meV}$  and  $n = 1$ , we exclude up to  $M < 1.7 \times 10^{-3} M_P$  (replacing  $M < 2.8 \times 10^{-3} M_P$ ); the gap to torsion pendulum constraints is fully closed for  $\Lambda > 6.0 \text{ meV}$  (replacing 5.1 meV), and for  $\mu = 0.1 \text{ meV}$ , we rule out  $\lambda < 0.8$  (replacing 1).

Except for these corrections, our conclusions remain unchanged. We replotted the exclusion plot (Fig. 3 in the original paper) with the corrected  $\alpha_{\text{anomaly}}$  value. Owing to the wide range of parameters considered in the plot, the correction is barely visible.

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**Fig. 1 | Original and Corrected Fig. 3.** Corrected Fig. 3 demonstrates constraints on screened chameleons (a, b) and symmetrons (c).