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# Inner core static tilt inferred from intradecadal oscillation in the Earth's rotation

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The presence of a static tilt between the inner core and mantle is an ongoing discussion encompassing the geodynamic state of the inner core. Here, we confirm an approximate 8.5 yr signal in polar motion is the inner core wobble (ICW), and find that the ICW is also contained in the length-of-day variations of the Earth's rotation. Based on the determined amplitudes of the ICW and its good phase consistency in both polar motion and the length-of-day variations, we infer that there must be a static tilt angle  $\theta$  between the inner core and the mantle of about  $0.17 \pm 0.03^{\circ}$ , most likely towards ~90°W relative to the mantle, which is two orders of magnitude lower than the 10° assumed in certain geodynamic research. This tilt is consistent with the assumption that the average density in the northwestern hemisphere of the inner core should be greater than that in the other regions. Further, the observed ICW period (8.5 ± 0.2 yr) suggests a  $0.52 \pm 0.05$  g/cm<sup>3</sup> density jump at the inner core boundary.

The oblate solid Earth consists mainly of a solid inner core, a liquid outer core, and a solid mantle with the same center of mass, which is reduced to a series of layered elliptical surfaces on which the density is constant in the classical model<sup>1,2</sup>. The current theories regarding the Earth's rotation involve the consideration of the mantle's elliptical surfaces of constant density, whose symmetry axes are aligned in the direction of rotation, and hydrostatic effects necessitate that the inner core's figure axis  $\Omega_{ic}$  (as defined in Appendix A of ref. 2) and rotation axis  $\Omega'_m$  are aligned with the mantle's figure or rotation axis  $\Omega_m$  (which are nearly identical due to centrifugal torque) in order to maintain equilibrium. The presence of random torques acting on the inner core results in a slight tilt and further excites a prograde rotation mode known as the inner core wobble (ICW), i.e., the inner core's figure axis  $\Omega_{ic}$  wobbles about its rotation axis  $\Omega'_m^{2-4}$  (also represents the direction of the lowest gravitational potential energy of the mantle-inner core system). The above 'tilt' between the  $\Omega_{ic}$  and  $\Omega'_m$  is a generally dynamic tilt, and in this case, the ICW theoretically appears only in the polar motion (PM) of the Earth's rotation<sup>2,3,5</sup>. In addition, the ICW period is very sensitive to the density jump  $\Delta \rho_{\rm ICB}$  at the inner core boundary (ICB;  $\Delta \rho_{ICB}$  is still not well constrained<sup>6</sup>), and based on the PREM model<sup>1</sup>, theory predicts that its period falls in the range of  $6.6-7.8 \text{ yr}^{3,7-9}$ . However, the elliptical surfaces of constant density within the heterogeneous mantle may exhibit random tilting around its rotation axis  $\Omega_m$  considering its solid properties, particularly with significant uncertainties at the core-mantle boundary (CMB). Consequently, the inner core's rotation axis  $\Omega'_{m}$ , which signifies the direction of the static equilibrium of the inner core (and corresponds to the lowest gravitational potential energy of the mantle-inner core system<sup>10</sup>), was previously believed to deviate from alignment with the mantle's axis  $\Omega_m$  and instead possess a tilt relative to it, and thereby the 'tilt' between the  $\Omega'_m$  and  $\Omega_m$  is called as static tilt. To explain the decadal oscillations in both the PM and the length-of-day variations ( $\Delta$ LOD) as the possible ICW, the inner core's rotation axis  $\Omega'_m$  was proposed to coincide with the dipole axis of the geomagnetic field (tilted 10° westwards from  $\Omega_m$ )<sup>11</sup>. Despite the absence of confirmed observations of the ICW<sup>12,13</sup> and the lack of universal acceptance of the excessive static tilt of 10°, the possibility of a static-tilted inner core remains, and further investigation has been conducted to explore the

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impact of a static tilt on the period of the ICW<sup>12</sup>. Theoretically, such a static tilt must affect some modes that are sensitive to the inner core. However, no relevant eigenfrequency deviation has been clearly detected in the core-sensitive normal modes of Earth's free oscillation<sup>14,15</sup>, which denotes that this static tilt is still uncertain, it may not exist or is quite small. Overall, a statically tilted inner core will be of great importance to some fundamental research about the Earth, such as the differential rotation of the inner core, the Earth's surface gravity changes, the seismic tomography of the deep Earth, and the Geodynamo theory<sup>6,12,14,16-20</sup>.

A statically tilted inner core will induce changes in the rotational normal modes of the Earth, with the ICW being the most sensitive. In the presence of a statically tilted inner core, the ICW will manifest not only in the PM but also in the  $\Delta LOD^{II}$ . By identifying a similar periodic signal in the  $\Delta LOD$  and establishing its correlation with the ICW identified in the PM, we can, in turn, ascertain the presence of a statically tilted inner core. Furthermore, the angle  $\theta$  of static tilt can be determined by comparing the corresponding amplitudes of the two signals (in the PM and in the  $\Delta LOD$ , respectively).

#### **Results and discussion**

# Inner core wobble in the polar motion and the length-of-day variations

In this study, we report the results from the  $\Delta$ LOD and PM time series. The chosen  $\Delta$ LOD time series is a yearly time series with a 1900–2020 time span. For the PM time series, the 1900-2020 EOPC01 time series with one-year sampling is used (x and y components). The pretreatments of the  $\Delta$ LOD and PM time series are shown in the Methods. Figure 1 shows the  $\Delta$ LOD and PM records used. For the periodic signals present in the PM and  $\Delta$ LOD, the consensus is that they are excited by the Earth's internal or external sources through the conversion of angular momentum<sup>21</sup>. Hence, we need to rule out the influence of external excitation sources before determining that a target signal is from the Earth's internal motion. There are three external excitation sources of the PM and  $\Delta$ LOD changes, the atmospheric, oceanic, and hydrological effects. Of these, the first two effects are the two main external excitation sources<sup>21-23</sup>; although hydrological effects will also excite the Earth's rotation changes, previous studies have proven that the hydrological effects have no significant contribution to the target 5.5-10 yr period band<sup>13,23</sup> and different hydrological models have clear deviations<sup>13,24</sup>. Hence, similar to previous studies<sup>13,22</sup>, we only consider



**Fig. 1** | **The used polar motion (PM) and length-of-day variations (\DeltaLOD) records.** The *x* component (**a**) and *y* component (**b**) of the observed PM records from 1900 to 2020 (gray curves) and the PM excited by the atmospheric and oceanic angular momentum (AAM + OAM, abbreviated as AO: black curve) from 1949 to 2020; **c** the observed  $\Delta$ LOD record from 1900 to 2020 (gray curves) and the  $\Delta$ LOD excited by the AAM + OAM (AO: black curve) from 1949 to 2020.

the atmospheric and oceanic effects. The PM and  $\Delta$ LOD excited by the atmospheric angular momentum (AAM) and oceanic angular momentum (OAM) are also shown in Fig. 1. Supplementary Fig. 1 (in the Supplementary Information) also shows all datasets used before combining.

Different from the oceanic tidal signals that have both prograde and retrograde components in PM<sup>25-27</sup>, the ICW is a prograde motion (the same as the Chandler wobble, i.e., the mantle wobble); in the complex spectra of the PM, a prograde/retrograde wobble only has a positive/negative frequency. The identification of the well-known Chandler wobble is based on this feature<sup>28-30</sup>. Therefore, as a prograde motion, the ICW only appears on the positive frequency axis of the PM spectrum and this is a distinguishing feature for identifying it. Based on the 1960-2017 PM record without removing the AO (AAM + OAM) effects, a previous study<sup>13</sup> used this feature to identify an ~8.7 yr signal for the ICW; here, we perform independent detection by using a longer record (1949-2020) and further consider the AO effects. Figure 2 shows the normalized AR-z spectra (see Methods) of the PM and  $\Delta$ LOD records in the 1949-2020 time span, in which the AO effects have been removed. Figure 2a shows four different harmonic signals (~5.9, ~7.3, ~8.5, and ~9-11 yr) in the positive frequency axis; only the ~8.5 yr signal has no corresponding spectral peak in the negative frequency axis (The AR-z method is meant for determining the presence of a signal and estimating its frequency, the amplitude of it contains no direct information about the actual complex amplitude of the detected signal). The corresponding Fourier spectra of the PMs (observed and AO excited) show similar findings (see Supplementary Fig. 2b). Among these harmonics, the ~5.9 yr signal has been suggested as the inner core oscillation coupled with torsional wave in the Earth's core but still remains controversial<sup>31-33</sup>; the ~7.3 yr signal can be interpreted as the Magneto-Coriolis eigenmode in the Earth's core based on a theoretical model<sup>34</sup>; the peak in the ~9–11 yr is possible from the ~11 yr Schwabe solar cycle. The adjacent ~13 yr signal (Fig. 2a) also has both positive and negative frequencies; similar period was also found in the geomagnetic dipole field<sup>35</sup>, but the underlying mechanism is still enigmatic: the ~18-23 vr spectral peak may be mainly caused by the 18.6 vr tidal signal and the ~22 yr Hale solar cycle or high-latitude MAC



Fig. 2 | Normalized AR-z spectra of the used polar motion (PM) and length-ofday variations ( $\Delta$ LOD) records (1949–2020). The AR-z spectra of the PM (a) and  $\Delta$ LOD (b) records in which the atmospheric and oceanic angular momentum effects were removed. Different colored rectangles indicate different periodic signals; the gray area represents 95% confidence intervals (C.I.).

(magnetic-Archimedes-Coriolis forces) wave<sup>36</sup> in the Earth's core; those periods are too long to be the ICW<sup>3,7-9</sup>. Therefore, the  $8.52 \pm 0.19$  yr signal is the only candidate for the ICW. Since no other mechanism has been proposed to account for such a prograde -8.5 yr motion, and the AO effects have been removed, we can conclude that the 8.5 yr signal is the ICW. In addition, the uncertainties for the estimates in this study were based on a bootstrap procedure<sup>37</sup>.

Comparing Fig. 2a and b, a finding is that the six periodic/quasiperiodic signals in the positive frequency axis of the PM spectra are also present in the spectrum of the  $\Delta$ LOD. This mainly benefits from the high-frequency resolution of the AR-z spectrum and its strong sensitivity to harmonic signals<sup>38</sup>; the Fourier spectra can only identify parts of those signals (see Supplementary Fig. 2). These consistencies deserve further attention, but we only focus on the 8.5 yr signal (the ICW signal). Figure 2b confirms that the ICW signal is also present in the  $\Delta$ LOD (with an 8.47 ± 0.32 yr period); this finding preliminarily suggests that there should be a static tilt between the inner core and the mantle. Given that the AO effects have no significant contribution to the target signal, we use the 1900–2020 PM and  $\Delta$ LOD records to extract the -8.5 yr signal to obtain higher resolutions. For simplicity, we directly use a cosine least-square fitting process.

#### Static tilt between the inner core and mantle

To further obtain the orientation of the static tilt angle  $\theta$  and its magnitude, we need to determine the fluctuation characteristics of the axial torque  $\Gamma_z$  ( $\propto d\Delta \text{LOD}/dt$ ; see Methods) exerted on the mantle. Hence, we directly fit the -8.5 yr signal from  $d\Delta \text{LOD}/dt$ ; the fitted results from  $\Delta \text{LOD}$  can be found in Fig. S3.

Figure 3 shows the fitted ICW from the  $d\Delta LOD/dt$  and the x and y components of the PM. Clearly, the ICW from the y-component is ahead of that from the x-component by  $-\pi/2$  (see green areas in Fig. 3); since the directions of x and y have a  $\pi/2$  angle difference in the equatorial plane, these findings are acceptable. The most important

point obtained from Fig. 3 is that, for the first time, we find that the ICW signals contained in the *y* component of the PM and  $d\Delta LOD/dt$  have almost synchronous phases; the extracted oscillations using a more complicated method (the normal time-frequency transform, NTFT<sup>39</sup>) show almost the same results (see Supplementary Fig. 3). This synchronicity is not a random phenomenon and at least demonstrates that the inner core tilts in a particular direction (see the possible scenario in Fig. 4). The axial torque  $\Gamma_z$  reaches its peak/trough only when the  $\Gamma_z$  is in the plane defined by the static tilted axis  $\Omega'_m$  and the rotation axis of the mantle  $\Omega_m$ ; hence, we can deduce that the inner core tilts should be along the -90°E–90°W direction.

Given the y component of the PM along the  $90^{\circ}$ W longitude, the phase synchronization in Fig. 3a, c indicates that the inner core is more



Fig. 3 | The least-square fitted -8.5 yr signals from different records. From (a) the first-order time derivative of length-of-day variations ( $d\Delta LOD/dt$ ); (b) the *x* component of the polar motion (PM), and (c) the *y* component of the PM. The time span used is 1900–2020, and the sampling interval is one year.



**Fig. 4** | **Schematic depiction of the tilted inner core wobble (ICW).** The figure axis of the inner core wobbles about its rotation axis tilted to the mantle in a small circular orbit (see blue shaded area).  $\Omega_m$  is the rotation axis of the mantle,  $\Omega'_m$  and  $\Omega_{ic}$  are respectively the rotation axis and dynamic figure axis of the inner core, and  $\theta$ 

is the static tilt angle between  $\Omega_m$  and  $\Omega'_m$ . The upper right subgraph shows how to get  $\theta$  from the mantle-inner core gravitational torque in the mantle's axial and equatorial directions  $|F_z^{\text{MICG}}|$  and  $|F_{eq}^{\text{MICG}}|$  (see also in Methods).

likely tilted in the 90°W direction, which is also similar to that suggested by previous studies<sup>11,12,40,41</sup>. In terms of the long-term dynamic conservation of the Earth's angular momentum, this static westwards tilt is consistent with the effect of the non-axisymmetric mass of the inner core, i.e., the western hemisphere (more specifically, the northwestern hemisphere) of the inner core should have greater average densities. To explain the asymmetry between the inner core's eastern and western hemispheres in seismological observations<sup>42,43</sup>, a previous dynamical model considers the crystallization and melting at the surface of the inner core and has similar suggestions<sup>44</sup>, i.e., the western hemisphere of the inner core is denser than its eastern hemisphere. Interestingly, a seismological study has suggested that the western hemisphere of the inner core may be relatively denser<sup>45</sup>, and a thicker compacting layer at the top of the inner core's western hemisphere was also suggested<sup>46</sup>; a more nuanced research found that the western zone is largely confined to the northern hemisphere<sup>47</sup>; those suggestions are generally consistent with the westwards statically tilted inner core that we found.

Here we can propose the following scenario as schematically depicted in Fig. 4: There is a static westward tilt  $\theta$  between the inner core and the mantle, resulting in the inner core exhibiting a wobbling motion around the tilted axis  $\Omega'_m$ . This wobbling motion leads to the exchange of angular momentum between the mantle and inner core in both the equatorial and axial directions of the mantle, consequently giving rise to the appearance of the ICW in both the PM and  $\Delta$ LOD. The inner core and mantle then obtain the maximum or minimum deviation at ~90°E – 90°W (i.e.,  $x \approx 0$ ) equatorial diameter; the torque on the inner core (equatorial plane of the inner core) thus has a maximum/minimum component on the axis of rotation when  $x \approx 0$ . At the same time, the  $\Delta LOD$  is also 0 due to a first derivative relationship with the exchange of the axial angular momentum (see Supplementary Fig. 3). Therefore, there will be a good phase consistency between the  $d\Delta LOD/dt$  and the y component of the PM for the ICW (confirmed in Fig. 3).

Next, we estimate the tilt  $\theta$ . The heterogeneous mantle results in a tilted rotation axis of the inner core  $\Omega'_m$ . When the inner core's figure axis  $\Omega_{ic}$  is aligned with  $\Omega'_{m}$ , the mantle-inner core system has the lowest mantle-inner core gravitational (MICG) potential energy (the equilibrium state). Any deviation of the inner core's figure axis  $\Omega_{ic}$ from the equilibrium state caused by the ICW will result in a MICG restoring torque  $\Gamma_{\text{MICG}}$ ; the torque  $\Gamma_{\text{MICG}}$  always brings the inner core back to the equilibrium state. Thus, the torque  $\Gamma_{MICG}$  is always in the plane (the blue plane in Fig. 4) perpendicular to the inner core's rotation axis  $\Omega'_m$ . The mean equatorial plane of the inner core is perpendicular to the axis  $\Omega'_{m}$ , so  $\Gamma_{\text{MICG}}$  is always in the mean equatorial plane of the inner core. This MICG torque can also be decomposed into the equatorial and axial torques exerted on the mantle due to the tilt of the inner core. If there is no static tilt between the inner core and mantle, the axial torque exerted on the mantle will be 0, and the entire gravitational torque will provide the equatorial torque exerted on the mantle.

Under the MICG coupling, the equatorial torque exerted on the mantle can be written as the following (see Methods):

$$\left| \boldsymbol{\Gamma}_{\text{eq}}^{\text{MICG}} \right| = (\boldsymbol{C}_m - \boldsymbol{A}_m) \boldsymbol{\Omega}_0^2 |\boldsymbol{\chi}_{\text{ICW}}(t)| \tag{1}$$

where  $\chi_{ICW}(t)$  is the excitation sequence of the ICW in the PM, and the observed ICW is almost the same as its excitation sequence due to the low frequency<sup>21,48</sup> (see Methods). Since the ICW identified in the PM has an amplitude of  $4.7 \pm 0.4$  mas, we can calculate  $|\Gamma_{eq}^{MICG}|$  to be  $(2.87 \pm 0.24) \times 10^{19} \,\text{N} \cdot \text{m}$ . According to the theorem of angular momentum, the axial torque exerted on the mantle is written as the following (see Methods):

$$\left|\boldsymbol{\Gamma}_{z}^{\text{MICG}}\right| = \frac{\Omega_{0}C_{m}}{\text{LOD}}\frac{d(\Delta\text{LOD}_{\text{ICW}})}{dt}$$
(2)

Substituting the amplitude of  $0.046 \pm 0.005 \text{ ms/yr}$  (corresponding to the amplitude of  $0.061 \pm 0.007 \text{ ms}$  of  $\Delta \text{LOD}_{\text{ICW}}$ ) and the period of 8.5 yr for the observed ICW signal in the  $\Delta \text{LOD}$  into Eq. (2), the axial torque  $|\boldsymbol{\Gamma}_z^{\text{MICG}}|$  is calculated to be  $(8.61 \pm 0.95) \times 10^{16} \text{ N} \cdot \text{m}$ ; this is only a small component of the equatorial torque of the inner core caused by the ICW due the inner core static tilt. Therefore, the static tilt angle  $\theta$ , or the angle between the axis about which the inner core wobbles and the rotation axis of the mantle, is calculated as  $\arctan(|\boldsymbol{\Gamma}_z^{\text{MICG}}|/|\boldsymbol{\Gamma}_{\text{eq}}^{\text{MICG}}|)$  (see Fig. 4) and equal to  $0.17 \pm 0.03^\circ$ ; this is much smaller than previous assumptions.

Our observed ICW period is slightly larger than the theoretical values (6.6–7.8 yr)<sup>3,7–9</sup>, but considering that even the generally accepted Chandler wobble observation period of prograde ~430 days is ~30 days longer than its theoretical periods<sup>28–30</sup>, free core nutation observation period of retrograde ~430 days is ~20 days shorter than its theoretical periods<sup>49,50</sup>, and that the density jump  $\Delta \rho_{\rm ICB}$  at the ICB was also poorly determined<sup>6,51</sup>, this deviation is accepted. Considering this newly determined period of the ICW, we can also invert the density jump  $\Delta \rho_{\rm ICB}$ . Taking the density profiles of the PREM model as a reference, we finally obtained  $\Delta \rho_{\rm ICB} = 0.52 \pm 0.05$  g/cm<sup>3</sup> (see Methods), which is smaller than that of the PREM model (0.598 g/cm<sup>3</sup>).

In summary, based on the Earth's rotation observations (PM and  $\Delta$ LOD), we experimentally confirmed for the first time that the 8.5 yr signal is the ICW. The evidence indicates that the inner core is tilted to the mantle along ~90°W, and the inverted tilt angle is  $0.17 \pm 0.03^{\circ}$ ; this static tilt angle means that the average density in the northwest hemisphere of the inner core should be greater. The larger observed period may also indicate that the eastwards differential rotation rate of the inner core should be much less than  $1^{\circ}$  per year<sup>12,16</sup>. Besides, the density jump of  $0.52 \pm 0.05$  g/cm<sup>3</sup> at the ICB is also inverted based on the observed ICW period. Undeniably, it is difficult for seismological observations to detect such inner core static tilt directly, but interestingly, the results from seismological studies showed that the western/northwestern hemisphere (or at least its top layer) of the inner core may be relatively denser<sup>42-47</sup>. These suggestions, although they have some uncertainties, are gualitatively consistent with our finding of a westwards-tilted inner core, and we suggest such consistency should be helpful to the inner core oscillation or differential rotation.

#### Methods

# Conservation of angular momentum of the mantle and inner core

Considering the mantle alone, the law of angular momentum can be rewritten as<sup>48</sup>:

$$\frac{d}{dt}\mathbf{H}_m + \boldsymbol{\Omega} \times \mathbf{H}_m = \boldsymbol{\Gamma}_m \tag{3}$$

where Earth's angular velocity  $\mathbf{\Omega} = \mathbf{\Omega}_0[m_1, m_2, 1+m_3]^T$ ; the mantle angular momentum is:

$$\mathbf{H}_m = \mathbf{I}_m \boldsymbol{\Omega} \tag{4}$$

The asymmetric part of the mantle mentioned above is insignificant relative to its axisymmetric part, and the mantle can still be approximately as axisymmetric in the calculation of torque for simplicity;  $I_m$  is the mantle moment of inertia tensor initially expressed in the principal axes:

$$\mathbf{I}_{m} = \begin{bmatrix} A_{m} & 0 & 0\\ 0 & A_{m} & 0\\ 0 & 0 & C_{m} \end{bmatrix}$$
(5)

Combining the above equations with the eigenfrequency of the free Euler wobble replaced by the Chandler wobble  $\sigma_{CW}$ , the equatorial

torque exerted on the mantle can be obtained by the observed PM

$$\frac{i}{\sigma_{\rm CW}} \frac{d\mathbf{m}}{dt} + \mathbf{m} = \boldsymbol{\chi}(t) = \frac{\boldsymbol{\Gamma}_{\rm eq}}{i(\boldsymbol{C}_m - \boldsymbol{A}_m)\boldsymbol{\Omega}_0^2}$$
(6)

where  $\Gamma_{eq} = \Gamma_1 + i\Gamma_2$  is the equatorial torque exerted on the mantle;  $\Omega_0 = 7.29212 \times 10^{-5} \text{ s}^{-1}$  is the mean (sidereal) rotation rate<sup>9</sup>;  $A_m = 7.0999 \times 10^{37} \text{ kg} \text{ m}^2$  and  $C_m = 7.1236 \times 10^{37} \text{ kg} \text{ m}^2$  are the equatorial and axial moments of inertia of the mantle, respectively;  $\mathbf{m} = m_1 + im_2 = x$ *-iy* is the observed PM;  $\chi(t) = \chi_1 + i\chi_2$  is its excitation function. The relation between an excitation function of complex frequency  $\sigma$  and the motion of the observed pole is  $\chi(t) = (1 - \sigma/\sigma_{CW})\mathbf{m}$  (in which  $\sigma_{CW} = \omega_{CW} + i\gamma_{CW}$ ,  $\omega_{CW} = 2\pi \cdot 0.843$  cpy and  $\gamma_{CW} = \omega_{CW}/2Q_{CW}$ ;  $Q_{CW} \approx 30-150)^{52,53}$ . Thus, there is little difference between excitation and observation in the low-frequency band.

Similarly, the axial torque exerted on the mantle can be obtained by the following:

$$\Omega_0 C_m \frac{dm_3}{dt} = \mathbf{\Gamma}_z \tag{7}$$

where  $m_3 = -\Delta \text{LOD/LOD}$ ; LOD = 86400 s. Combining with Eqs. (1), (2), (6) and (7), we can directly infer the static tilted angle of the inner core from the ICW signal in the  $\Delta \text{LOD}$  and PM, which is impossible in related previous studies<sup>11,12</sup>.

#### Stabilized AR-z spectrum

A real discrete time series with the length of N equally spaced samples, which contains M harmonics, is written as (which satisfies the AR relation<sup>54</sup>):

$$x(n) = \sum_{j=1}^{M} \left[ \mathbf{A}_{j} \exp(in\sigma_{j}) + \mathbf{A}_{j}^{*} \exp(-in\sigma_{j}^{*}) \right], n = 1, 2, 3, \dots, N$$
(8)

where  $\mathbf{A}_j = A_j \exp(i\phi_j)/2$  is the complex amplitude  $(A_j \text{ and } \phi_j \text{ are the amplitude and initial phase})$  and  $\sigma_j = \omega_j + i\alpha_j$  is the complex frequency of a given harmonic  $(\omega_j \text{ and } \alpha_j \text{ are the angular frequency and decay rate})$ . By using a frequency-domain AR method, the complex frequency  $\sigma_j$  can be estimated<sup>38</sup>. A Lorentzien power spectrum in the complex *z* plane can be formed as follows<sup>38</sup>:

$$P(\sigma_i) = \frac{1}{|\exp(\tilde{\sigma}_i) - \exp(i\sigma_i)|^2}, i = 1, 2, \dots, N$$
(9)

where  $\tilde{\sigma}_i$  and  $\sigma_i$  are the estimated and referred complex frequencies, respectively. For the specific execution of the stabilized AR-z spectrum, please see the Supplementary Text (in the Supplementary Information).

#### Constraint for the density jump at the inner core boundary

The frequency of the ICW can be written as follows (in cpsd: cycle per solar day)<sup>3</sup>:

$$\sigma_{\rm ICW} = [\alpha_3(1+\alpha_g)(e_s + S_{34}^g + S_{34}^p)]/(1+K^{\rm ICB})$$
(10)

where the elastic compliances  $S_{34}^g = -1.812 \times 10^{-6}$ ,  $S_{34}^p = -2.686 \times 10^{-4}$  and  $K^{4CB}$  is a dimensionless coupling constant<sup>55</sup> and Real( $K^{4CB}$ ) =  $1.11 \times 10^{-3}$ ;  $\alpha_3$  and  $\alpha_g$  have the following forms:

$$\begin{cases} \alpha_3 = 1 - (A'e'/A_s e_s)\alpha_g \\ \alpha_g = \frac{3G}{a_s^2 \Omega^2} \left( \left[ \frac{5\rho}{3\rho_f} + 1 \right] A'e' - A_s e_s \right) - 1 \end{cases}$$
(11)

where  $\rho_f$  is the fluid density just outside the ICB,  $\bar{\rho}$  is the mean density of the inner core,  $A_s$  and  $e_s$  are the equatorial moment of inertia and the

dynamical ellipticity of the inner core, respectively, and A' and e' have similar definitions but for a body of the inner core radius with the constant mass of that of the fluid core at the ICB<sup>2</sup>.

$$\begin{cases} A_{s} = \frac{8\pi}{3} \int_{0}^{a_{s}} \rho(r) \{r^{4} - \frac{1}{15} \frac{d[\varepsilon(r) \cdot r^{5}]}{dr} \} dr \\ A'e' = \frac{8\pi}{15} \rho_{f} a_{s}^{5} \varepsilon_{s} \\ A'(1 + \frac{e'}{3}) = \frac{8\pi}{15} \rho_{f} a_{s}^{5} \end{cases}$$
(12)

in which  $\varepsilon(r)$  is the geometrical ellipticity of the Earth and  $a_s$  is the inner core radius. Taking the PREM model as a reference because it is the generally accepted model, underlying the conservation of the whole Earth's mass and angular momentum, we can modify the density of the outer core (based on the related expression given in PREM) and hence change the inner core density profile to obtain the observed ICW period. When the observed 8.5 yr period is obtained, the corresponding  $\Delta \rho_{\rm ICB}$  is the one we recommend using.

#### Datasets and preprocessing

The PM observations were obtained from the EOPC01 dataset (1861/ 01-1889/12 with 0.1 yr sampling and 1900/01-2019/12 with 0.05 yr sampling); the  $\Delta$ LOD record was combined with a long-term dataset<sup>56</sup> (1623/06-2008/06 with 1 yr sampling from IERS; EOPC01) and the EOPC04 ΔLOD record<sup>57</sup> (1962/01-2019/12 with 1-day sampling); the AAM (1949/01-2019/12, sampling at 6 h) record was from the Special Bureau for the Atmosphere<sup>58-60</sup>. The AAM was calculated from NCEP/ NCAR reanalyses archived on pressure surfaces, and the inverted barometer (IB) pressure term was chosen as the mass term. The OAM record was obtained from the Special Bureau for the Oceans' datasets: ECCO 50 yr<sup>61</sup> (1949/01-2003/01, sampling at 10 days) and ECCO kf080i<sup>62</sup> (1993/01-2020/3, sampling at 1 day). Those datasets are shown in Supplementary Fig. 1. To standardize the sampling intervals of the records, we down-sampled all records to 1 yr, and to avoid aliasing effects in this down-sampling process, a low-pass filter (with a cut-off frequency  $f_c = 0.5$  cpy) was used prior to downsampling.

Note that although the theoretical amplitudes of the 8.85 yr and 9.3 yr zonal tides are quite small (only ~2 µs for the 9.3 yr tide and <1 µs for the 8.85 yr tide) and far less than the background noise level of the  $\Delta$ LOD time series, they were removed from this  $\Delta$ LOD record based on a given model<sup>63</sup> to avoid the effects of some well-known signals on the target ~8 yr period band. The  $d\Delta$ LOD/dt was obtained by a classical discrete numerical derivation algorithm, i.e.,  $d\Delta$ LOD( $t_i$ )/  $dt = [\Delta$ LOD( $t_{i+1}$ ) –  $\Delta$ LOD( $t_i$ )]/ $\Delta t$ .

### Data availability

All associated source data including Earth's rotation and external excitations have been deposited in the Figshare database and can be accessed at https://doi.org/10.6084/m9.figshare.22820525; those data can be used to reproduce the results shown in Figs. 1–3.

#### Code availability

The code of the AR-z spectrum has been uploaded to https://doi.org/ 10.1029/2018JB015890. It is also available from the corresponding author H.D upon request.

#### References

- Dziewonski, A. M. & Anderson, D. L. Preliminary reference Earth model. Phys. Earth Planet. Inter. 25, 297–356 (1981).
- Mathews, P. M., Buffett, B. A., Herring, T. A. & Shapiro, I. I. Forced nutations of the Earth: influence of inner core dynamics: 1. Theory. J. Geophys. Res. Solid Earth 96, 8219–8242 (1991).
- Dumberry, M. Influence of elastic deformations on the inner core wobble. *Geophys. J. Int.* 178, 57–64 (2009).

### https://doi.org/10.1038/s41467-023-43894-9

- Rochester, M. G., Crossley, D. J. & Zhang, Y. L. A new description of Earth's wobble modes using Clairaut coordinates: 1. Theory. Geo-
- phys. J. Int. 198, 1848–1877 (2014).
  Guo, J. Y., Greiner-Mai, H. & Ballani, L. A spectral search for the inner core wobble in Earth's polar motion. J. Geophys. Res. Solid Earth 110, B10402 (2005).
- 6. Tkalčić, H. The Earth's Inner Core: Revealed by Observational Seismology (Cambridge Univ. Press, 2017).
- Rochester, M. G. & Crossley, D. J. Earth's long-period wobbles: a Lagrangean description of the Liouville equations. *Geophys. J. Int.* 176, 40–62 (2009).
- Crossley, D. J. & Rochester, M. G. A new description of Earth's wobble modes using Clairaut coordinates 2: Results and inferences on the core mode spectrum. *Geophys. J. Int.* **198**, 1890–1905 (2014).
- Mathews, P. M., Buffett, B. A., Herring, T. A. & Shapiro, I. I. Forced nutations of the Earth: Influence of inner core dynamics: 2. Numerical results and comparisons. J. Geophys. Res. Solid Earth 96, 8243–8257 (1991).
- Chao, B. F. Dynamics of axial torsional libration under the mantleinner core gravitational interaction. J. Geophys. Res. Solid Earth 122, 560–571 (2017).
- Greiner-Mai, H., Jochmann, H. & Barthelmes, F. Influence of possible inner-core motions on the polar motion and the gravity field. *Phys. Earth Planet. Inter.* **117**, 81–93 (2000).
- Guo, J. Y. & Ning, J. S. Influence of inner core rotation and obliquity on the inner core wobble and the free inner core nutation. *Geophys. Res. Lett.* 29, 45–1 (2002).
- Ding, H., Pan, Y., Xu, X. Y., Shen, W. & Li, M. Application of the AR-z spectrum to polar motion: a possible first detection of the inner core wobble and its implications for the density of Earth's core. *Geophys. Res. Lett.* 46, 13765–13774 (2019).
- 14. Deuss, A. Heterogeneity and anisotropy of Earth's inner core. *Annu. Rev. Earth Planet. Sci.* **42**, 103–126 (2014).
- Ding, H. & Chao, B. F. Data stacking methods for isolation of normalmode singlets of Earth's free oscillation: extensions, comparisons, and applications. J. Geophys. Res. 120, 5034–5050 (2015).
- Song, X. & Richards, P. G. Seismological evidence for differential rotation of the Earth's inner core. *Nature* 382, 221–224 (1996).
- 17. Dumberry, M. Decadal variations in gravity caused by a tilt of the inner core. *Geophys. J. Int.* **172**, 921–933 (2008).
- Tkalčić, H. Complex inner core of the Earth: the last frontier of global seismology. *Rev. Geophys.* 53, 59–94 (2015).
- Triana, S. A. et al. Core eigenmodes and their impact on the Earth's rotation. Surv. Geophys. 43, 107–148 (2022).
- Wang, W. & Vidale, J. E. Earth's inner core rotation, 1971 to 1974, illuminated by inner-core scattered waves. *Earth Planet. Sci. Lett.* 577, 117214 (2022).
- Gross, R. in *Treatise on Geophysics* 1st edn. (ed. Schubert, G.) 239–294 (Elsevier, 2007).
- Holme, R. & De Viron, O. Characterization and implications of intradecadal variations in length of day. *Nature* 499, 202–204 (2013).
- Chen, J., Wilson, C. R., Kuang, W. & Chao, B. F. Interannual oscillations in earth rotation. J. Geophys. Res. Solid Earth 124, 13404–13414 (2019).
- 24. Jin, S., Chambers, D. P. & Tapley, B. D. Hydrological and oceanic effects on polar motion from GRACE and models. *J. Geophys. Res. Solid Earth* **115**, B02403 (2010).
- Herring, T. A. & Dong, D. Measurement of diurnal and semidiurnal rotational variations and tidal parameters of Earth. J. Geophys. Res. Solid Earth 99, 18051–18071 (1994).
- Chao, B. F. & Ray, R. D. Oceanic tidal angular momentum and Earth's rotation variations. *Prog. Oceano.* 40, 399–421 (1997).

- 27. Gross, R. S. The excitation of the Chandler wobble. *Geophys. Res.* Lett. **27**, 2329–2332 (2000).
- 28. Chandler, S. C. On the variation of latitude, II. *Astron. J.* **11**, 65–70 (1891).
- 29. Newcomb, S. On the periodic variation of latitude, and the observations with the Washington primevertical transit. *Astron. J.* **11**, 81–82 (1891).
- Colombo, G. & Shapiro, I. I. Theoretical model for the Chandler wobble. *Nature* 217, 156–157 (1968).
- Buffett, B. A. Gravitational oscillations in the length of day. *Geophys.* Res. Lett. 23, 2279–2282 (1996).
- Buffett, B. A. & Mound, J. E. A Green's function for the excitation of torsional oscillations in the Earth's core. J. Geophys. Res. Solid Earth 110, B08104 (2005).
- Gillet, N., Jault, D., Canet, E. & Fournier, A. Fast torsional waves and strong magnetic field within the Earth's core. *Nature* 465, 74–77 (2010).
- Gillet, N. et al. Satellite magnetic data reveal interannual waves in Earth's core. Proc. Natl Acad. Sci. USA 119, e2115258119 (2022).
- Jin, R. S. & Thomas, D. M. Spectral line similarity in the geomagnetic dipole field variations and length of day fluctuations. *J. Geophys. Res.* 82, 828–834 (1977).
- 36. Nicolas, Q. & Buffett, B. Excitation of high-latitude MAC waves in Earth's core. *Geophys. J. Int.* **233**, 1961–1973 (2023).
- Efron, B. & Tibshirani, R. Bootstrap methods for standard errors, confidence intervals, and other measures of statistical accuracy. *Stat. Sci.* 1, 54–75 (1986).
- Ding, H. & Chao, B. F. Application of stabilized AR-z spectrum in harmonic analysis for geophysics. J. Geophys. Res. Solid Earth 123, 8249–8259 (2018).
- Liu, L., Hsu, H. & Grafarend, E. W. Normal Morlet wavelet transform and its application to the Earth's polar motion. J. Geophys. Res. Solid Earth 112, B08401 (2007).
- 40. Greiner-Mai, H. Possible relations between the rotational axis of the earth's inner core and the magnetic dipole axis. *Astron. Nachr.* **318**, 63–71 (1997).
- Szeto, A. M. K. & Smylie, D. E. Motions of the inner core and mantle coupled via mutual gravitation: regular precessional modes. *Phys. Earth Planet. Inter.* 54, 38–49 (1989).
- Tanaka, S. & Hamaguchi, H. Degree one heterogeneity and hemispherical variation of anisotropy in the inner core from PKP (BC)–PKP (DF) times. J. Geophys. Res. Solid Earth **102**, 2925–2938 (1997).
- Creager, K. C. Large-scale variations in inner core anisotropy. J. Geophys. Res. Solid Earth 104, 23127–23139 (1999).
- Alboussière, T., Deguen, R. & Melzani, M. Melting-induced stratification above the Earth's inner core due to convective translation. *Nature* 466, 744–747 (2010).
- 45. Monnereau, M., Calvet, M., Margerin, L. & Souriau, A. Lopsided growth of Earth's inner core. *Science* **328**, 1014–1017 (2010).
- Zhang, Y., Nelson, P., Dygert, N. & Lin, J. F. Fe alloy slurry and a compacting cumulate pile across Earth's inner-core boundary. J. Geophys. Res. Solid Earth 124, 10954–10967 (2019).
- Brett, H., Hawkins, R., Waszek, L., Lythgoe, K. & Deuss, A. 3D transdimensional seismic tomography of the inner core. *Earth Planet. Sci. Lett.* **593**, 117688 (2022). (2022).
- Kuang, W., Chao, B. F. & Chen, J. Decadal polar motion of the Earth excited by the convective outer core from geodynamo simulations. J. Geophys. Res. Solid Earth 122, 8459–8473 (2017).
- 49. Zhang, H. & Shen, W. Core–mantle topographic coupling: a parametric approach and implications for the formulation of a triaxial three-layered Earth rotation. *Geophys. J. Int.* **225**, 2060–2074 (2021).

# Article

- 50. Getino, J. & Ferrándiz, J. M. Forced nutations of a two-layer Earth model. *Mon. Not. R. Astron. Soc.* **322**, 785–799 (2001).
- 51. Masters, G. & Gubbins, D. On the resolution of density within the Earth. *Phys. Earth Planet. Inter.* **140**, 159–167 (2003).
- 52. Furuya, M. & Chao, B. F. Estimation of period and Q of the Chandler wobble. *Geophys. J. Int.* **127**, 693–702 (1996).
- 53. An, Y. & Ding, H. Revisiting the period and quality factor of the Chandler wobble and its possible geomagnetic jerk excitation. *Geod. Geodynamics* **13**, 427–434 (2022).
- Chao, B. F. On the use of maximum entropy/autoregressive spectrum in harmonic analysis of time series. *Pure Appl. Geophys.* 134, 303–311 (1990).
- Mathews, P. M., Herring, T. A. & Buffett, B. A. Modeling of nutation and precession: new nutation series for nonrigid Earth and insights into the Earth's interior. *J. Geophys. Res. Solid Earth* **107**, 2068 (2002).
- Stephenson, F. R. & Morrison, L. V. Non-tidal changes in the length of the day: 700 BC to AD 1982. *Geophys. Surv.* 7, 201–210 (1985).
- 57. Bizouard, C. Continental drift of the rotation pole: observation and theory. *IOP Conf. Ser.: Mater. Sci. Eng.* **468**, 012005 (2018).
- Kalnay, E. et al. The NCEP/NCAR 40-year reanalysis project. B. Am. Meteorol. Soc. 77, 437–472 (1996).
- Barnes, R. T. H., Hide, R., White, A. A. & Wilson, C. A. Atmospheric angular momentum fluctuations, length-of-day changes and polar motion. *Pro. R. Soc. Lond. A* 387, 31–73 (1983).
- Salstein, D. A., & Rosen, R. D. Global momentum and energy signals from reanalysis systems. 8th Conference on Climate Variations, 2–7 February 1997, Long Beach, California, American Meteorological Society, Boston, Massachusetts, 344–348 (1997).
- Gross, R. S., Fukumori, I. & Menemenlis, D. Atmospheric and oceanic excitation of decadal-scale Earth orientation variations. J. Geophys. Res. Solid Earth 110, B09405 (2005).
- Gross, R. S. An improved empirical model for the effect of longperiod ocean tides on polar motion. J. Geod. 83, 635–644 (2009).
- 63. Ray, R. D. & Erofeeva, S. Y. Long-period tidal variations in the length of day. J. Geophys. Res. Solid Earth **119**, 1498–1509 (2014).

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# **Author contributions**

H.D. and W.J. supervised the project. Y.A. and H.D. conceived the idea and designed the experiments. Y.A. conceived the data analysis and wrote the first draft. H.D. prepared the figures, assisted in the overall conceptualization and interpretation of results, and contributed substantially to the writing of the manuscript. Z.C., W.J. and W.S. contributed to discussions of the data and analysis. All authors participated in the writing of the manuscript.

### **Competing interests**

The authors declare no competing interests.

# **Additional information**

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