various sulphur and other compounds and the physical and optical effects of adsorbed gases such as H₂S and SO₂ thereon need to be studied as possible strong contributors to Io's spectral reflectance and other surface properties.

We thank Dennis L. Matson for helpful conversations, and

Received 1 June; accepted 12 July 1979.

- Fanale, F. P., Johnson, T. V. & Matson, D. L. Science 186, 922-925 (1974).
 Nash, D. B. & Fanale, F. P. Icarus 31, 40-80 (1977).
- 3
- Wamsteker, W., Kroes, R. L. & Fountain, J. A. Icarus 23, 417-424 (1974). Nelson, R. M. & Hapke, B. W. Icarus 36, 304-329 (1978).
- Harris, D. L. in The Solar System Vol. III. Planets and Satellites (eds Kuiper, G. P. & 5. Middlehurst, B. W.) 272-342 (University of Chicago Press, 1961).
- Johnson, T. V. Icarus 14, 94-111 (1971).
- Johnson, T. V. & McCord, T. B. Astrophys. J. 169, 589-593 (1971). 8. Pollack, J. B. et al. Icarus 36, 271-303 (1978).
- 9. Cruikshank, D. P. Icarus (in the press).
- 10. Morrison, D., Cruikshank, D. P. & Burns, J. A. in Planetary Satellites (ed. Burns, J. A.) 3-17 (University of Arizona Press, 1977).

George Rossman and Lois Taylor for technical assistance. This work represents the results of one phase of research carried out at the California Institute of Technology JPL under contract NAS 7-100, sponsored by NASA. R.M.N. is an NRC resident research associate.

- 11. Johnson, T. V. & Pilcher, C. B. in Planetary Satellites (ed. Burns, J. A.) 232-268 (University of Arizona Press, 1977).
- 12. Morrison, D. & Morrison, N. D. in Planetary Satellites (ed. Burns, J. A.) 363-378 (University of Arizona Press, 1977).
- 13
- Caldwell, J. Icarus 25, 384-396 (1975). Matson, D. L., Johnson, T. V. & Fanale, F. P. Astrophys. J. Lett. 192, L43-L46 (1974). 14. 15. Herzberg, G. Infrared and Raman Spectroscopy of Polyatomic Molecules (Van Nostrand,
- New York, 1945). 16. Nakamoto, K. Infrared Spectra of Inorganic and Coordination Compounds (Wiley, New
- York, 1963). 17. Eischen, R. & Pliskin, W. Advances in Catalysis Vol. 10, 1-56 (Academic, New York, 1958).
- 18. Lebofsky, L. Icarus 25, 205-217 (1975).
- 19. Nash, D. B. J. geophys. Res. 71, 2517-2534 (1966).

Spectral evidence for SO₂ frost or adsorbate on Io's surface

William D. Smythe^{*†}, Robert M. Nelson^{*}, & Douglas B. Nash^{*}

*Jet Propulsion Laboratory, California Institute of Technology, Pasadena, California 91103 [†]Institute of Geophysics and Planetary Physics, University of California at Los Angeles, Los Angeles, California 90024

The infrared reflection spectrum of Io matches that from an SO_2 frost grown in the laboratory.

IT has been recently suggested that adsorbed SO₂ and/or H₂S on the surface of a spectroscopically benign substrate can explain the IR spectrum of Jupiter's satellite, Io^1 . SO₂ gas has been identified as a major constituent of the volcanic gas plumes recently discovered by the Voyager 1 spacecraft². We have measured the reflection spectrum of SO₂ frost grown under low pressure at 140 K in the laboratory and find that it provides an excellent match to Io's spectrum in the 1.0-4.5 µm range.

The apparatus used and experimental methods have been described by Kieffer³ and Smythe⁴ respectively. Figure 1 compares the spectral geometric albedo of Io5.6 and our laboratory spectrum. Clearly SO₂ frost can explain the Io absorption bands at 1.35, 2.55, 2.80, 3.80, 3.90, 3.95, and 4.08 µm. Other candidate materials such as sulphates, nitrates and carbonates³ do not provide such a unique match to these bands.

Although SO₂ frost may be unstable on Io, except at the poles or nightside (see ref. 7), adsorbed SO₂ should be present on the surface particulates due to the extensive volcanic outgassing. The absorption bands of adsorbed and condensed species are known to be similar in wavelength⁸. Slight shifts in band positions and depths may reflect areal or temporal variation in frost/adsorbate or condensate/substrate contribution ratios to a multicomponent spectrum. Mixtures of alkali sulphides, free sulphur and adsorbates of SO₂ or H₂S have been suggested to explain Io's overall UV-VIS-IR reflectance¹. We conclude here that SO₂ as a frost or an adsorbate is the primary contributor to Io's IR spectral reflectance.

We thank Dennis L. Matson and Robert Carlson for helpful conversations, Hugh Kieffer for use of his laboratory at UCLA, and Larry Pleskot for assistance. We thank Dale Cruikshank for a revised Io spectra. This represents the results of one phase of research carried out at the JPL under NASA grant NAS 7-100. R.M.N. is an NRC resident research associate.

Received 8 June; accepted 17 July 1979.

- 1. Nash. D. B. & Nelson, R. M. Nature 280, 763-766 (1979).
- 2. Hanel, R. et al. Press Release (American Geophysical Union meeting, Washington, D.C., 5 May 1979).
- Kieffer, H. H. Appl. Opt. 8, 2497-2500 (1969).
- Smythe, W. D. thesis, Univ. California, Los Angeles (1979).
- 5. Pollack, J. B. et al. Icarus 36, 271-303 (1978).
- 6. Cruikshank, D. P. Icarus (in the press) Lebofsky, L. A. Icarus 25, 205-217 (1975)
- 8. Eischens, R. P. & Pliskin, W. A. Adv. Catalysis 10, 1-56 (1958).



Fig. 1 *a*, Spectral geometric albedo of Jupiter's satellite Io, 1.0-4.5 μ m based on previous studies^{5,6}, normalised as in ref. 1. O, ●, Data from ref. 5; □, data from ref. 6. b, Spectral reflectance of SO₂ frost grown in the laboratory. The 3.1 µm H₂O band (stippled area) is due to contaminant water.