Todhunter; but to make it effective our teachers must be possessed of ordinary common-sense. So long as this is absent, all the elaborate and scientifically improved editions of Euclid's "Elements" in the world will not produce the much-to-be-desired change. Let the teacher go through any edition of the first book of Euclid's "Elements" in a common-sense manner with his pupils, and he will find that, instead of the apathy and general disgust exhibited by them when undergoing the ordinary process of Euclidian cram, there will be a general air of brightness, interest, and intelligent appreciation. H.

The Yorkshire College, Leeds, November 25.

A Brilliant Meteor.

WHILE at my observatory to-night, at 9.37 p.m., I saw the largest and brightest meteor I have seen since November 1880. It became visible near ν Eridani, and disappeared near α Leporis. The colour was a bright greenish blue, and the brightness was twice or three times Venus at greatest brilliancy. It cast a distinct shadow. J. COCKBURN.

St. Boswells, N.B., November 23.

STAR DISTANCES.¹

THE festal offering contributed by Prof. Oudemans to the Pulkowa celebration is an especially appropriate one. The incidents of the long parallax-campaign can scarcely be recapitulated without recalling, in connection with the name of Friedrich Struve, the *quorum pars magna fui* of Æneas. He it was who, in Sir John Herschel's opinion (Memoirs R. Astronomical Society, vol. xii. p. 442), made the first real impression upon the problem by showing that not one of twenty-seven circumpolar stars discussed in 1819-21 could possibly have an annual parallax amounting to half a second of arc. Thenceforward, astronomers knew what they had to expect. Sanguine hopes of meeting comfortably large, and properly periodical residuals among ordinary observations, were checked, if not extinguished. The changes of stellar position reproducing, according to the laws of perspective, the movement of the earth in its orbit, were perceived to be on a scale so minute that their satisfactory disclosure lay, for the moment, beyond the range of what was feasible. Success in the enterprise, it was evident, was conditional upon the employment of more perfect instruments than had heretofore been available with a precision and vigilance of which the very idea was absent from all but a few prescient minds. Sir William Herschel seemed to have anticipated the conjuncture when he declared in 1782 the case to be "by no means desperate," although stellar parallax should fall short of a single second (*Phil. Trans.*, vol. lxxii. p. 83). The memorable "triple event," by which, almost simultaneously, at the Cape, at Königsberg, and at Pulkowa, his confidence was justified, is familiar to all readers of astronomical history. Its significance may be estimated from Bessel's admission that, until the yearly oscillations of 61 Cygni emerged from his measures in 1838, he was completely in the dark as to whether stellar parallax was to be reckoned by tenths or by thousandths of a second (*Astr. Nach.*, No. 385).

The value to students of Prof. Oudemans' synoptical view of what has since then been achieved in this direction can hardly be overstated. Not only does he record every individual result worth considering, but the tabulated particulars enable a fair judgment to be formed as to the value of each. There are, indeed, one or two cases in which a note of warning might with advantage have been added. Thus, Dr. Brünnow's small

parallax for 85 Pegasi, to say the least, requires confirmation. A perfect *equability* in the mode of observing is essential in such delicate operations; but the Dunsink astronomer was himself conscious of, and noted with his usual care, a slight change, as the series flowed on, in his habit of "bisecting" the large star (*Dunsink Observations*, vol. ii. p. 38). The distance of this interesting binary system can hence scarcely be regarded as even approximately known.

Still less reliable, though for different reasons, are Johnson's measures of Castor, and Captain Jacob's of α Herculis. The parallax assigned to the latter star of $0''.062$ relative to its fifth magnitude companion cannot be other than illusory, since the pair, as evidenced by a small, but well-ascertained common proper motion, are physically connected, and must therefore be at virtually the same distance from the earth.

Forty-nine stars, all save one measured within the last sixty years, are included in Prof Oudemans' list. The exception deserves particular mention. Samuel Molyneux erected at his house in Kew Green in 1725, a zenith sector by Graham, with which he began, in combination with Bradley, a set of observations for parallax on γ Draconis. The same star had, in the previous century, been similarly experimented upon by Robert Hooke with something of a dubious success. The well-known eventual issue of Molyneux's observations was Bradley's discovery of the aberration of light; but they included besides an element of true parallactic change, brought out by Dr. Auwers's discussion in 1869,¹ after it had lain concealed among them for 142 years. The eye and hand must indeed have been faithful thus to record an ebb and flow of change profoundly submerged, at that comparatively remote epoch, in the reigning confusion between the real and the apparent places of the heavenly bodies.

A light-journey of sixty-five years (parallax = $0''.05$) may be considered the present limit of really measurable stellar distance. Forty of the forty-nine objects so far investigated lie—most of them certainly, a few only probably—within it. Forty stars can thus be located with some definiteness in space—forty among, say, forty millions! The disproportion between our knowledge on the point and our ignorance is so exorbitant that general conclusions seem discredited beforehand, and negative ones at any rate can have no weight whatever. Nevertheless, one remark at least is fully warranted by the evidence.

It is this, that the largest stars are not always those nearest to the earth. For to the narrow category of stars at ascertained distances belong no less than seven invisible to the naked eye, one of them in closer vicinity to us than Sirius, all than Capella, Vega, Arcturus, or Canopus. A cursory view might almost suggest—irrespective of geometrical possibilities—that stellar brightness had nothing whatever to do with remoteness. The legitimate and certain conclusion to be derived from the facts, however, is that the disparities of stellar light-power are enormous. A farthing rushlight is not more insignificant compared with the electric arc than a faint compared with a potent sun. Sirius emits 6400 times as much light as a ninth magnitude star north of Charles's Wain (Argelander-Oeltzen 11,677); our own sun falls nearly as far short of the radiative strength of Arcturus. Inequalities of the same order between the members of revolving systems emphasize this result. Sirius shines like four thousand of its own companions; and the movements of other stars are perhaps swayed by almost totally obscure bodies.

The inference that the apparent lustre of individual stars tells us nothing as regards their distance was already

¹ "Übersicht der in den letzten 60 Jahren ausgeführten Bestimmungen von Fixstern parallaxen." Von J. A. C. Oudemans. Eine Festschrift zum 50-jährigen Jubiläum der Sternwarte zu Pulkowa. *Astronomische Nachrichten*, Nos. 2915-16.

² *Monatsberichte*, Berlin, 1869. p. 630. The result places γ Draconis at a distance of $35\frac{1}{2}$ light-years, but with a very large "probable error" (parallax = $0''.092 \pm 0''.070$).