

► concerns, but say that they are reserving judgement about whether to censor the papers until the US government provides details of how it will allow genuine researchers to obtain redacted information.

Steinbruner is frustrated that no such mechanism exists. About five years ago, he and his collaborators studied a scenario similar to the H5N1 transmissibility studies, and realized that the government would need to construct a system to disseminate such information to a limited, vetted set of users (J. Steinbruner *et al.* *Controlling Dangerous Pathogens: A Prototype Protective Oversight System*; Center for International and Security Studies at Maryland, Univ. Maryland, 2007). He says they notified the NSABB of their findings, to no effect.

Keim says that working out how to distribute sensitive information is “not our job”, adding that it may fall to the US government’s Department of Health and Human Services (DHHS), possibly with guidance from the White House Office of Science and Technology Policy.

Other biosecurity experts say that the NSABB should have been involved in reviewing the H5N1 work earlier. “The time when action is needed is at the grant-application level,” says Laura Kahn, a security expert at Princeton University in New Jersey. But the NSABB can only evaluate a project when the DHHS explicitly asks it to do so. Keim and several other board members hope that in the future they will be asked to weigh in earlier.

“All along, we hoped to assist the US government in putting in place a much more far-reaching, proactive oversight system,” says David Relman, a microbiologist at Stanford University in Palo Alto, California, and a member of the board since it began. But “everything of this sort comes at some cost, not just in terms of work and burden and time, but also dollars and bodies”.

For security experts such as Steinbruner, the best solution is to replace the board, or to supplement its role, with a mandatory biosecurity oversight system. Even some scientists who have argued strongly in favour of self-regulation say that they have been disappointed by the NSABB’s performance. “I wanted to see them do more,” says Atlas.

He may get his wish. In the course of its deliberations over the H5N1 papers, the NSABB became aware of additional work on H5N1 transmissibility that was nearing publication. Keim says the board is now considering whether to recommend a voluntary moratorium on the publication of such work until the community can discuss further precautions to prevent misuse. He expects the board to vote on this in the next few weeks, and adds: “It is time for us to have a broad and global discussion.” ■



Holes in aluminium plates channel light from individual galaxies to calculate how fast each is receding.

#### COSMOLOGY

# Survey tunes in to dark energy

*Sonic yardstick offers a measure of Universe's expansion.*

BY ERIC HAND

David Schlegel’s tool for exploring dark energy, one of nature’s biggest mysteries, is deceptively simple. It is an aluminium plate the size of a manhole cover — or rather, 2,200 of them, each with a specific pattern of holes drilled to match the arrangement of galaxies in a particular section of the sky. Every plate is used once, for an hour, at the prime focus of the 2.5-metre telescope at Apache Point Observatory in New Mexico. When the telescope is pointing at the correct spot, light from each galaxy streams through its corresponding hole. The light is then broken up into its constituent wavelengths and used to clock how fast each galaxy is being carried away from us by the relentless expansion of space.

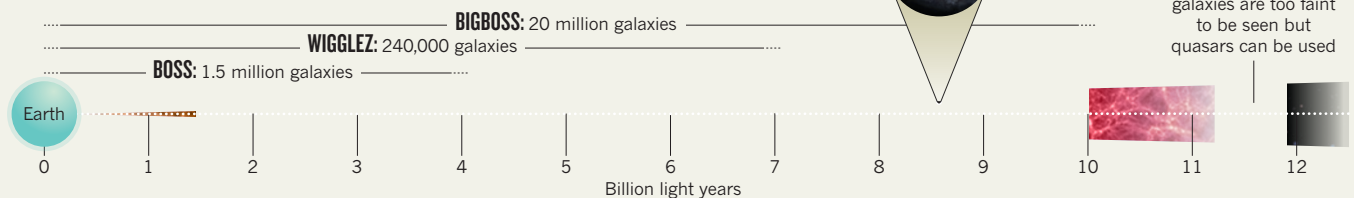
The study, which began in 2009, will ultimately gather data from 1.5 million galaxies. Its goal is to measure dark energy — a

phenomenon thought to be driving the Universe to expand at an ever-increasing rate — and discern whether its influence has remained constant or has varied slightly across billions of years of cosmic history. “The more galaxies we get, the better we do,” says Schlegel, an astronomer at Lawrence Berkeley National Laboratory (LBNL) in California and principal investigator of the massive undertaking, known as the Baryon Oscillation Spectroscopic Survey (BOSS).

But BOSS — and the method behind it — is already coming into its own. On 11 January, Schlegel and his colleagues will unveil their initial findings — based on data from 470,000 galaxies — at a meeting of the American Astronomical Society in Austin, Texas. Those data have given them a glimpse of cosmic structure by showing where galaxies clump together, like the crests of giant waves. The structure is a relic of a much younger and smaller Universe, in which acoustic waves reverberated through

## SIZING UP COSMIC RIPPLES

Astronomers map individual galaxies to determine the spacing between large-scale structures — a relic of acoustic waves generated in the early Universe. As the Universe expands, so do the spaces, revealing the influence of dark energy.



the dense, hot plasma that had not yet cooled enough to form stars and galaxies. These waves, called baryon acoustic oscillations (BAOs), pushed matter into regions of high and low concentrations with fairly even spacing — a pattern that evolved, in later epochs, into the giant sheets and filaments of galaxies that are the largest structures in the Universe.

First detected in 2005, the regular spacing between these structures creates a natural cosmic yardstick<sup>1</sup>. In today's Universe, that spacing has grown to about 150 megaparsecs (500 million light years). By detecting deviations from this yardstick, BOSS offers the tightest constraint yet on dark energy's influence (see 'Sizing up cosmic ripples').

Surveys of powerful stellar explosions known as type Ia supernovae yielded the first clues to dark energy in 1998. These supernovae are all thought to reach roughly the same peak brightness, allowing them to be used as 'standard candles' to determine the distances to their host galaxies. When these data are combined with measurements of how fast the supernovae are receding, they reveal that the expansion of the Universe is accelerating, rather than slowing down under gravity. One explanation is that dark energy is the 'cosmological constant', a repulsive outward pressure that is innate to the vacuum of space. BOSS is expected to narrow the uncertainty bounds on this model to a few per cent. That's "a factor of a few better than anything we have today," says Schlegel.

Chris Blake, an astronomer at Swinburne University of Technology in Melbourne, Australia, says that "BOSS will be the state of the art for the next five years, there's no doubt about that". Blake was co-investigator on the WiggleZ survey, which studied BAOs by mapping almost 240,000 galaxies using the 3.9-metre Anglo-Australian Telescope at Siding Spring Observatory in Australia. That was enough for the WiggleZ team to announce last year that it had verified, independently of supernova observations, that galaxies are behaving as though dark energy exists<sup>2,3</sup>.

But none of the probes of dark energy has been able to say with any certainty that dark energy has behaved as a constant throughout the Universe's history. To address this key question, the BOSS team has set its sights on BigBOSS, a survey that would sample a much larger volume of space, observing 20 million

galaxies up to 3.1 billion parsecs away, along with 4 million quasars — the luminous cores of even more distant galaxies. This would allow it to trace the influence of dark energy across most of cosmic history and to discern whether it really has remained constant.

In place of BOSS's aluminium plates, which are laboriously positioned and connected by hand, BigBOSS will rely on an automated system that can manoeuvre fibre tips into precisely the positions needed to gather the light of distant galaxies. The collaboration has submitted a US\$70-million proposal to the National Optical Astronomy Observatory to

upgrade the 4-metre Mayall telescope on Kitt Peak in Arizona and begin five years of observing in 2018.

Schlegel says that BigBOSS would exceed the BAO-measurement capabilities of much more

expensive space missions aiming to study dark energy, such as NASA's proposed \$1.6-billion Wide-Field Infrared Survey Telescope. "BigBOSS scoops a lot of what we thought we had to go to space for," he says.

The rise of BAO experiments such as BOSS and WiggleZ comes just as the statistical power of supernova studies is flagging as a result of uncertainties about how consistent the explosions really are. Moreover, the BAO method has more to offer — in addition to measuring the separation between vast clusters of galaxies, it can help to tease out how gravity affects galaxies within the clusters. "We can see the flows of those galaxies into those clusters," says Blake.

This could rule out the possibility that dark

energy is something other than a cosmological constant — for example, a change in general relativity at large scales. "It could be that there's something funny going on with gravity," says Josh Frieman, an astronomer at Fermilab in Batavia, Illinois.

Frieman's own tool for finding out is weak lensing, which measures how the gravity of large clumps of matter subtly distorts the shapes of more distant galaxies. Frieman is leading the Dark Energy Survey, which later this year will begin taking pictures of 300 million galaxies with a 570-megapixel camera installed at the 4-metre Blanco telescope at the Cerro Tololo Inter-American Observatory in Chile. But atmospheric blurring, which has hampered earlier efforts to map weak lensing with ground-based telescopes, could prove to be a challenge.

Because the BAO method requires only a galaxy's spectrum and not a high-resolution image, atmospheric blurring isn't as much of a problem. And the same data that BigBOSS would gather to identify the BAO effect on large scales can also be used to observe what gravity is doing around individual clusters of galaxies.

Saul Perlmutter, the LBNL astronomer who shared a Nobel prize for the discovery of cosmic acceleration, says that even if BigBOSS doesn't reveal the nature of the effect, it is bound to deliver some unexpected pay-offs. He points out that his own supernova survey — one of the two that led to the initial discovery — was originally expected to measure the deceleration of the Universe due to gravity. "We're on a roll," he says. "We have to keep pushing." ■

1. Eisenstein, D. J. *et al. Astrophys. J.* **633**, 560–574 (2005).
2. Blake, C. *et al.* preprint at <http://arxiv.org/abs/1105.2862> (2011).
3. Blake, C. *et al.* preprint at <http://arxiv.org/abs/1104.2948> (2011).

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